

Numerical Methods

(Additional Notes from Talks by PFH)

SPH

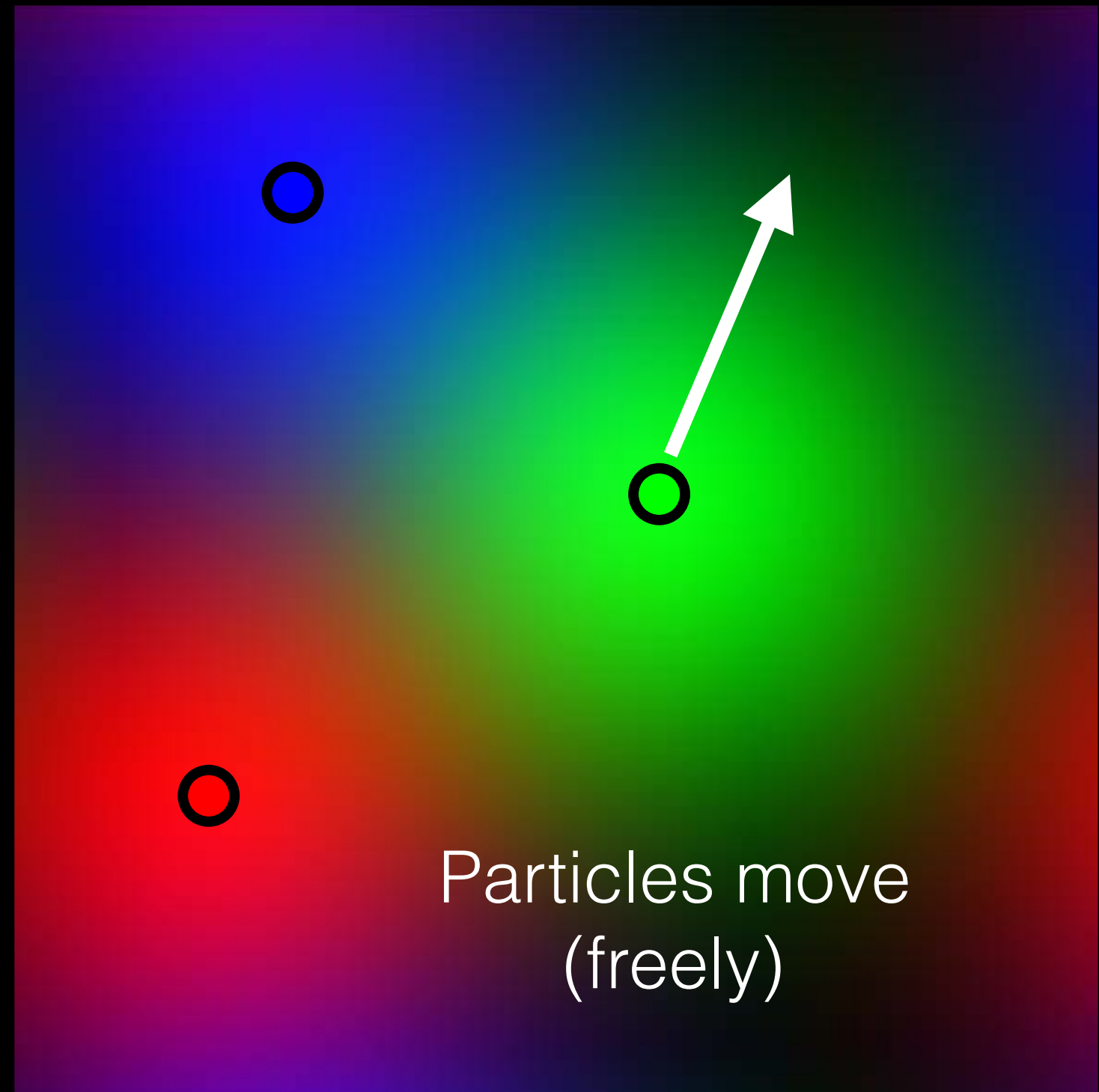
Challenge:

POPULAR METHODS FOR
HYDRODYNAMICS HAVE PROBLEMS

Lucy 77, Gingold & Monaghan 77
Reviews by: Springel 11, Price 12

Smoothed-Particle Hydrodynamics

- Lagrangian, adaptive,
simple, conservative

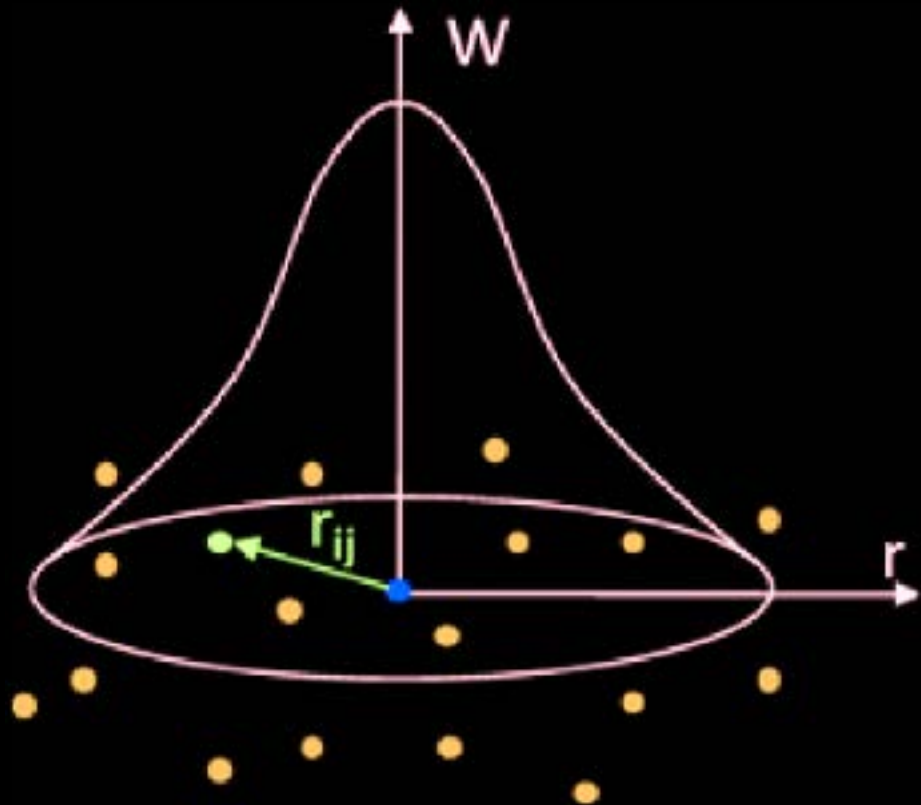


Challenge:

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HYDRODYNAMICS HAVE PROBLEMS

Smoothed-Particle Hydrodynamics

- No volume partition: point-like particles smoothed into fields [ok in “continuum limit”]



(average)

$$\frac{\partial \mathbf{v}_i}{\partial t} = \frac{\nabla P}{\rho} \Big|_{\mathbf{x}_i}$$

- Solve EOM at particle locations (stabilize with artificial diffusion)

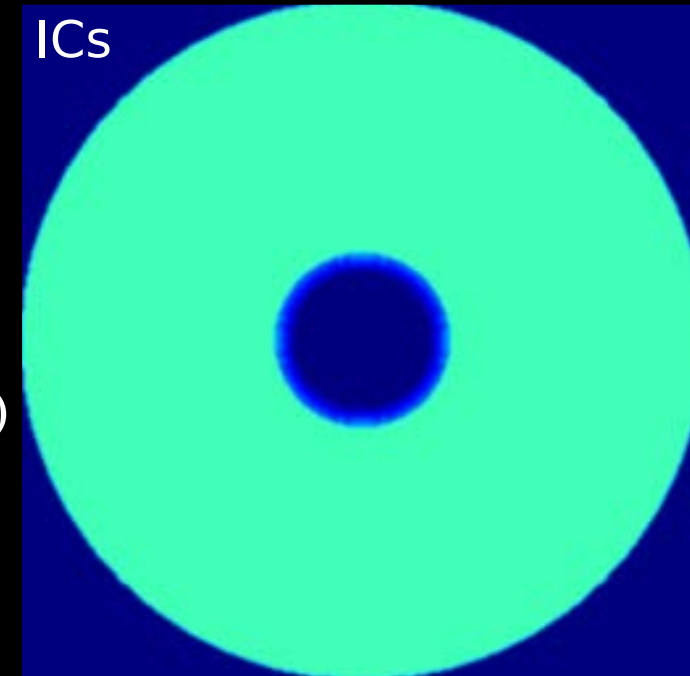
Challenge:

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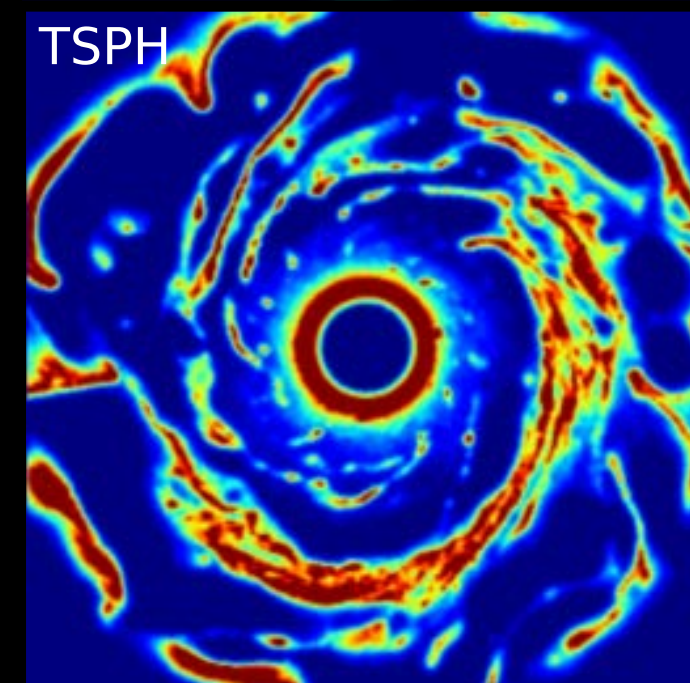
Smoothed-Particle Hydrodynamics

- Lagrangian, adaptive, simple, conservative
- Artificial diffusion terms:
 - excess diffusion, viscosity

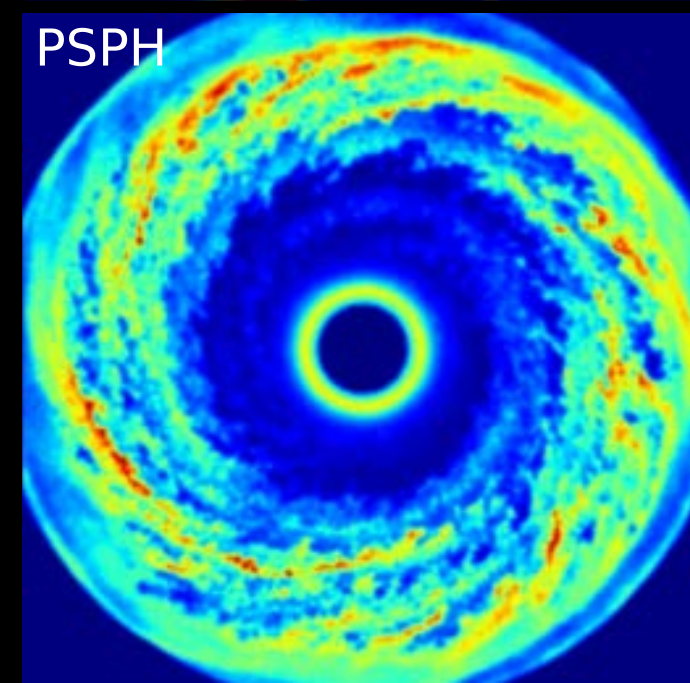
Keplerian disk/ring
(should conserve ICs)



“old” SPH
(Springel 02)
(after 20 orbits)



“new” SPH
(Hopkins 13)



Challenge:

POPULAR METHODS FOR HYDRODYNAMICS HAVE PROBLEMS

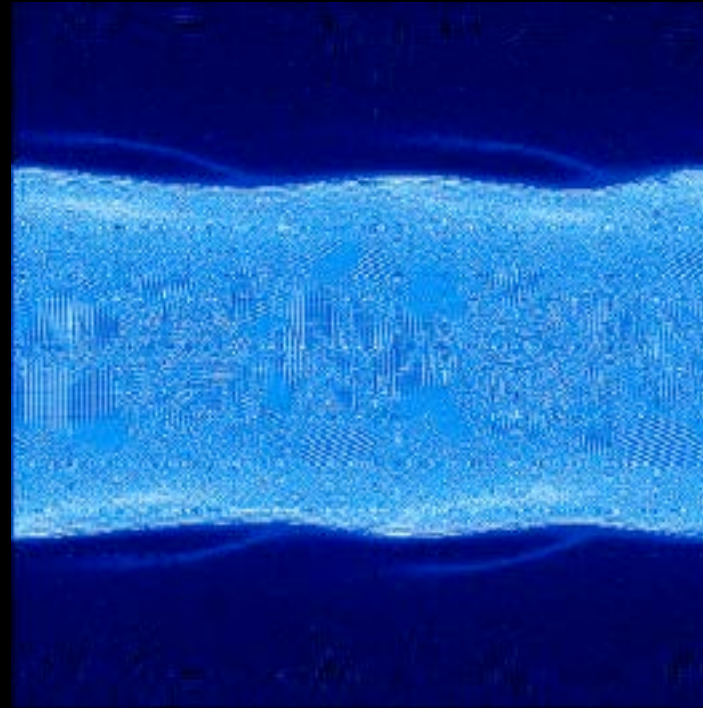
- **“Traditional SPH”**

- GADGET/(old)GASOLINE
- ~32 neighbors (cubic spline)
- constant artificial viscosity
- “density” formulation

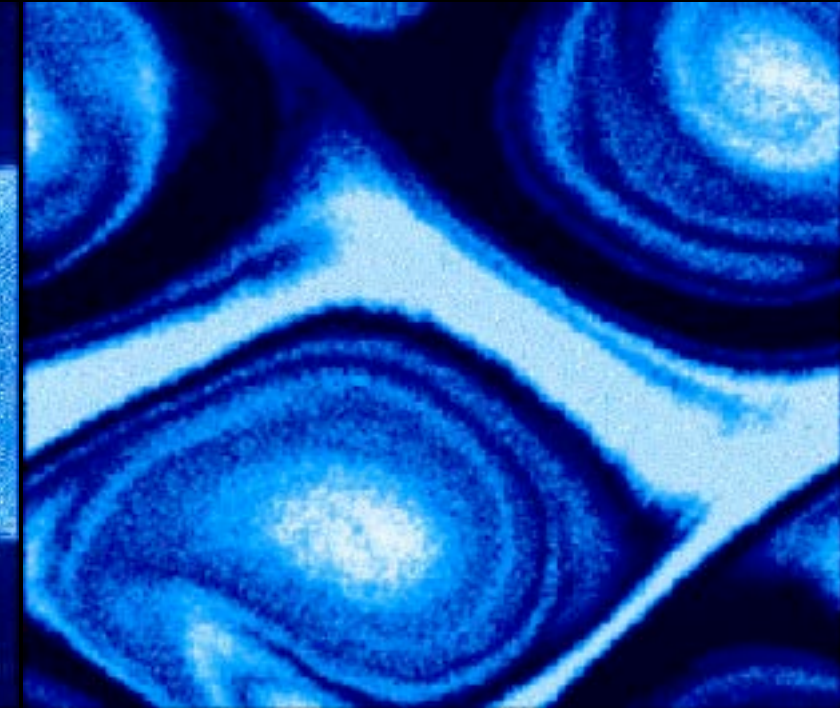
- **“Modern SPH”**

- P-SPH/SPHS/PHANTOM
- ~128-500 neighbors (alt. kernels)
(many people: Read, Dehnen)
- high-order switches
(Cullen+Dehnen)
- “pressure” formulation
(Hopkins, Saitoh+Makino)
- artificial diffusion for entropy
(Price, Wadsley)

Kelvin-Helmholtz Instabilities

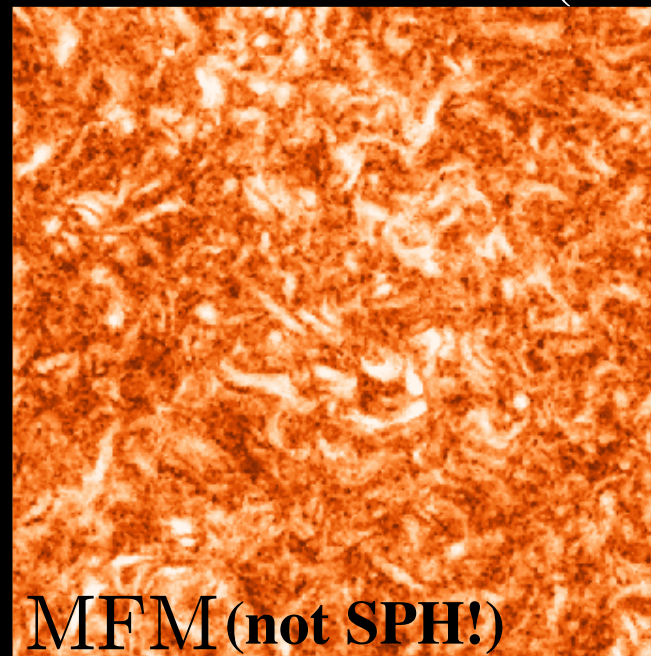


“old” SPH
(Springel 02)

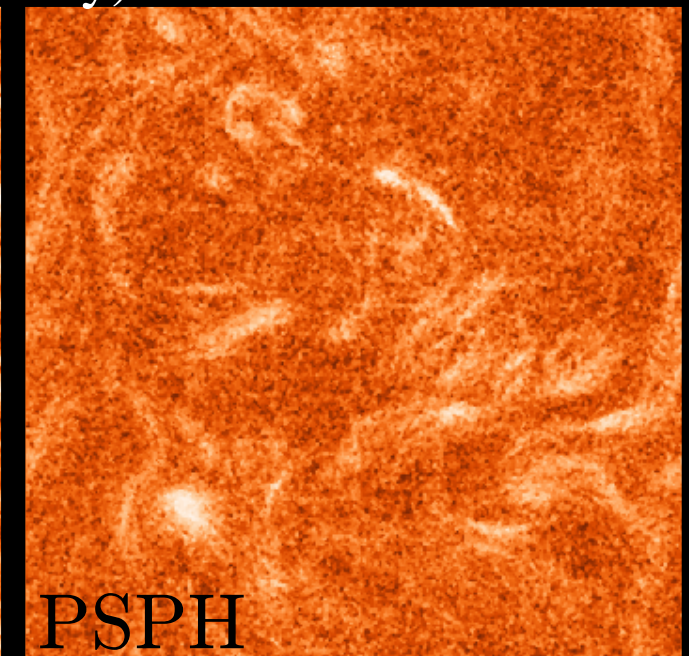


“new” SPH (PSPH)
(Hopkins '13): >>100 neighbors

Sub-sonic turbulence (vorticity)

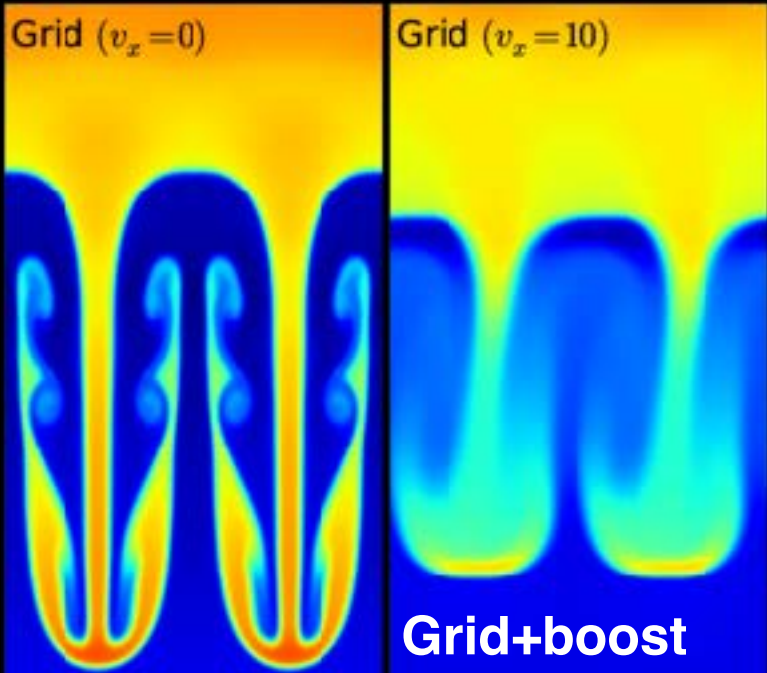
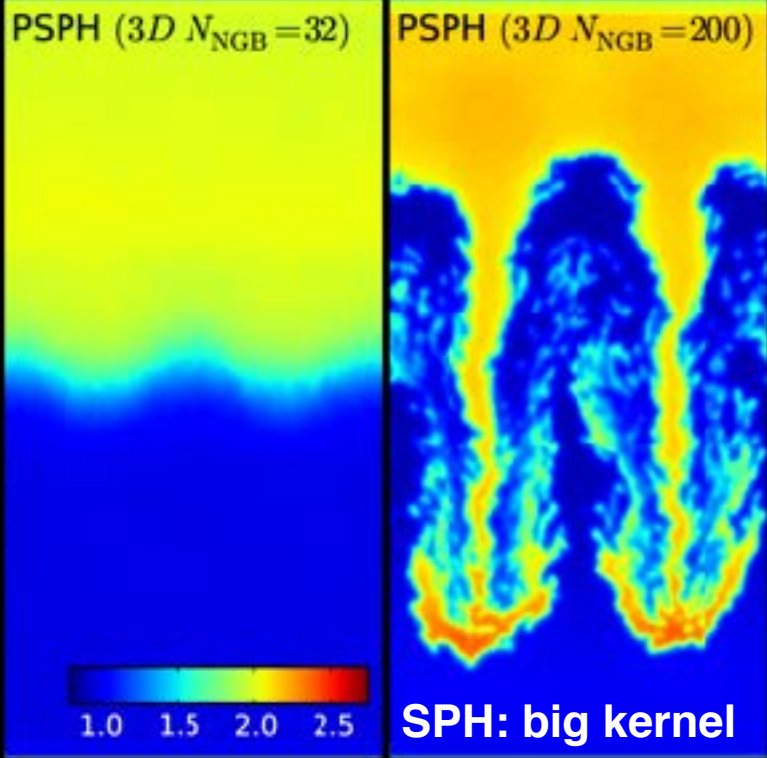
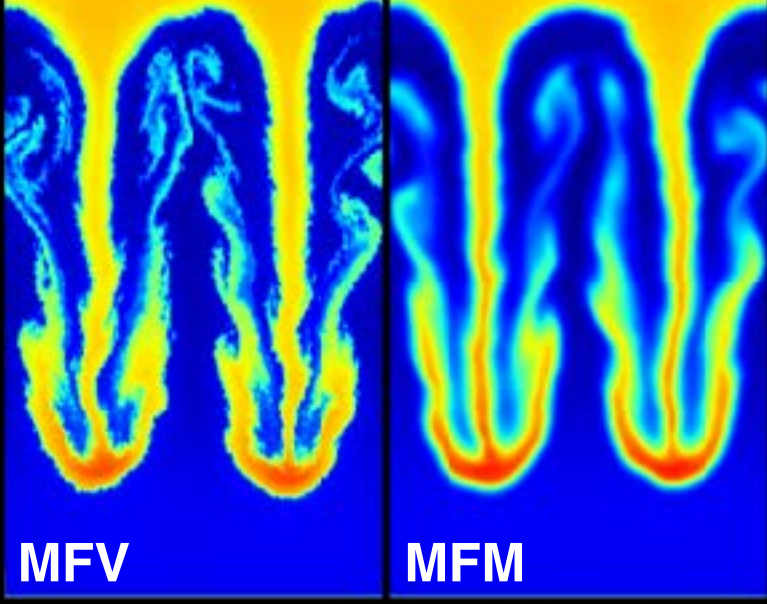
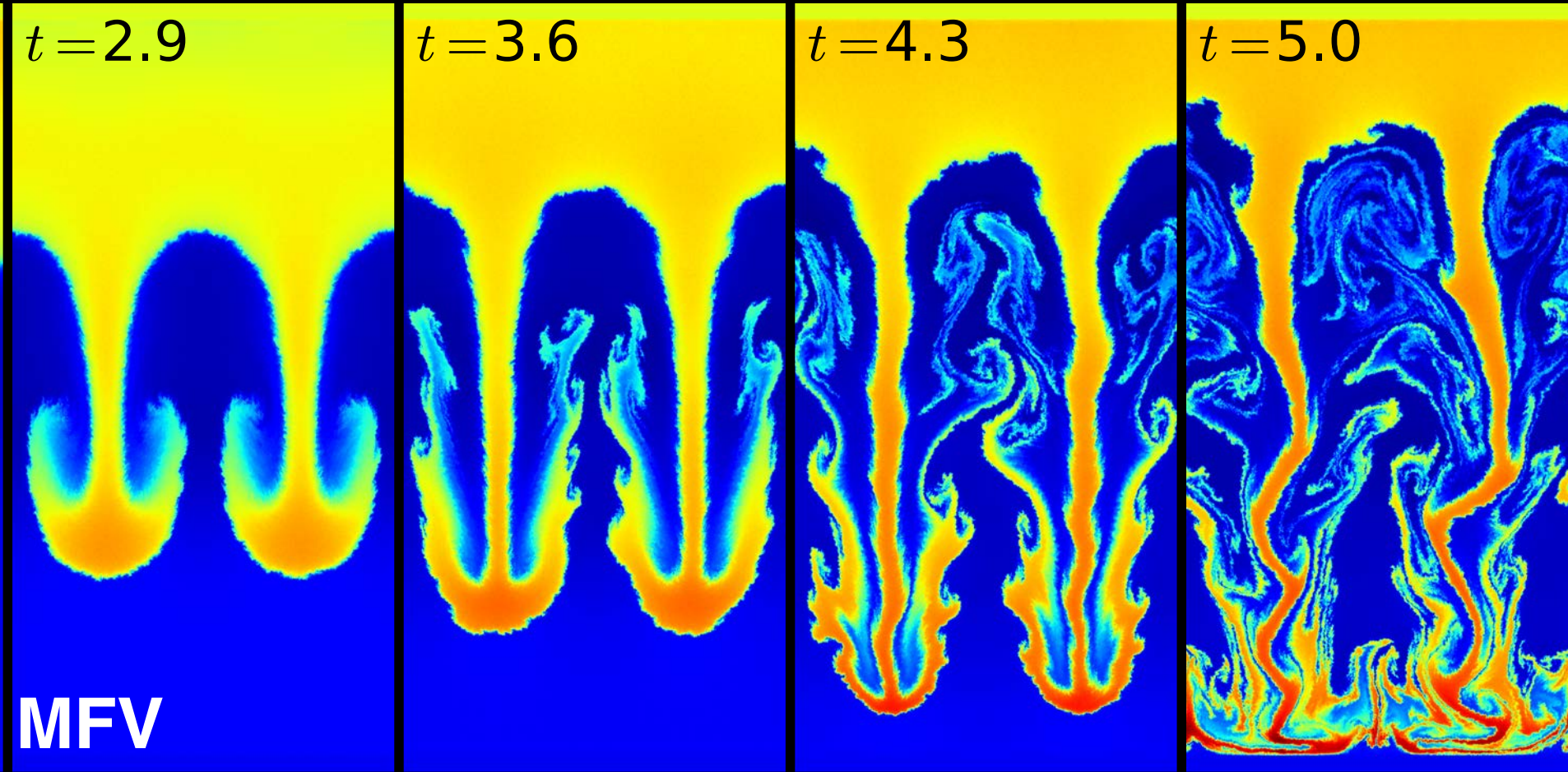


MFM (not SPH!)

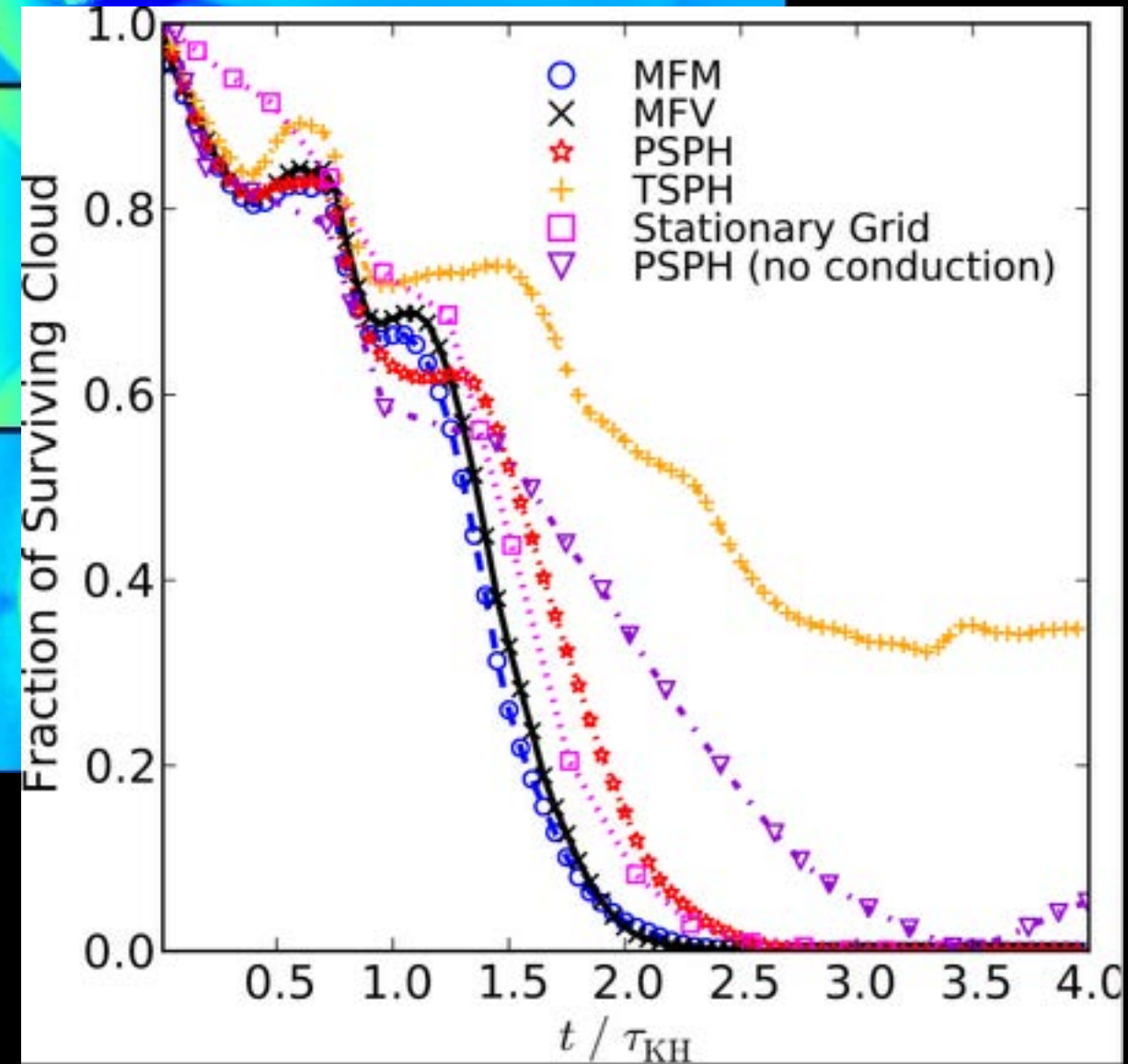
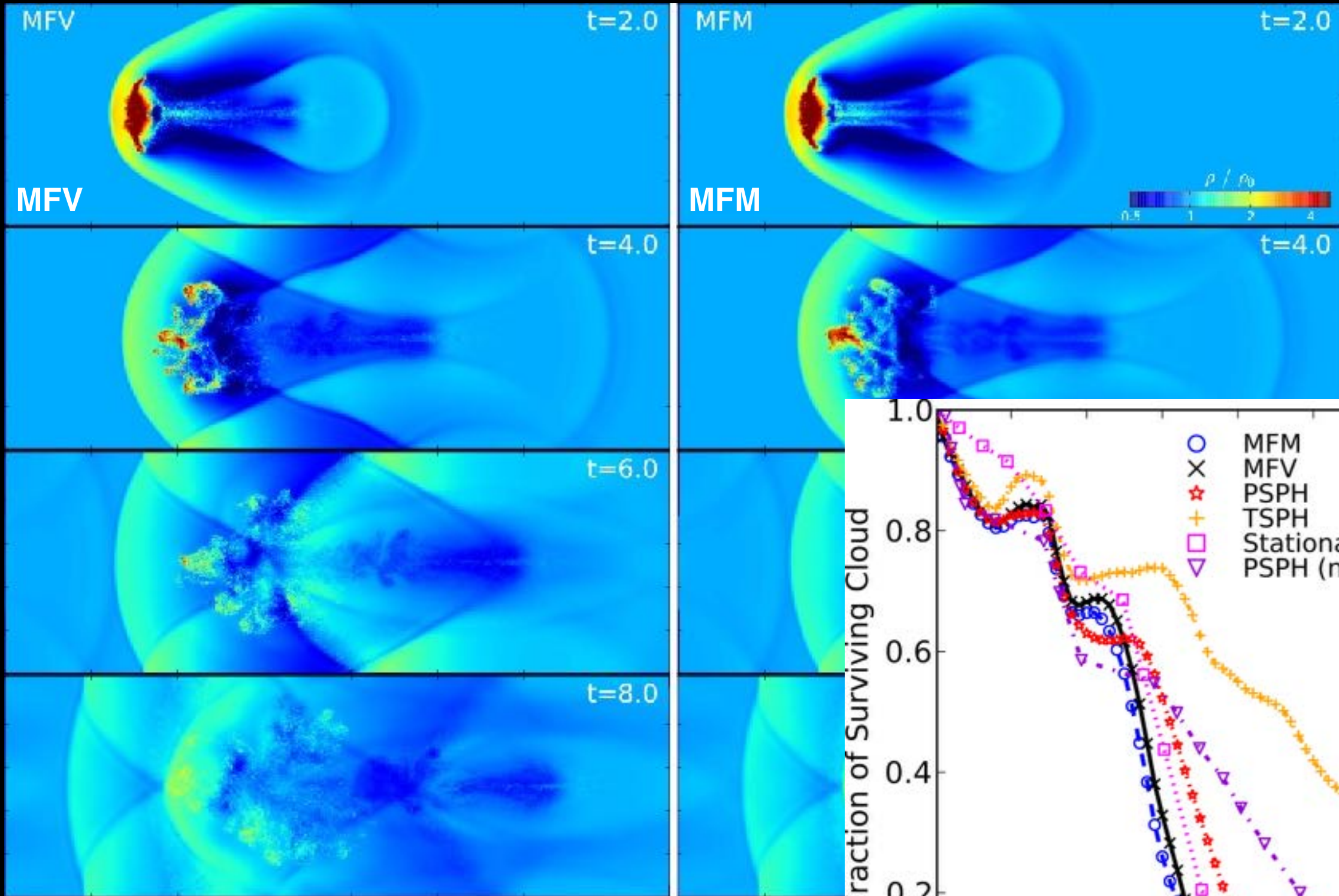


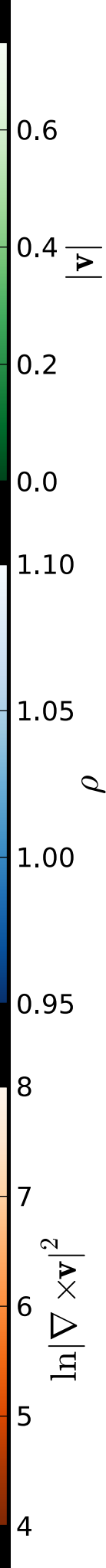
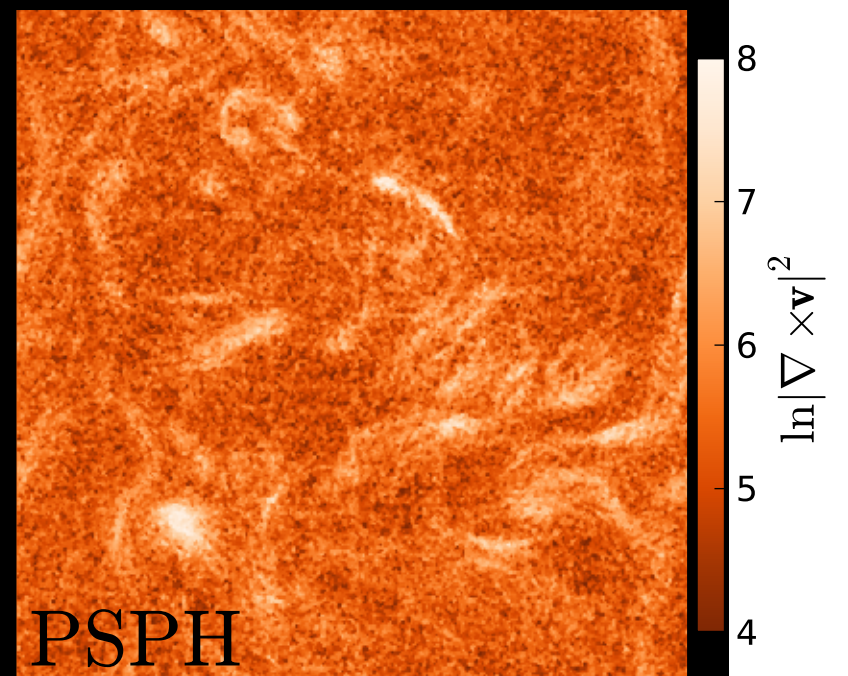
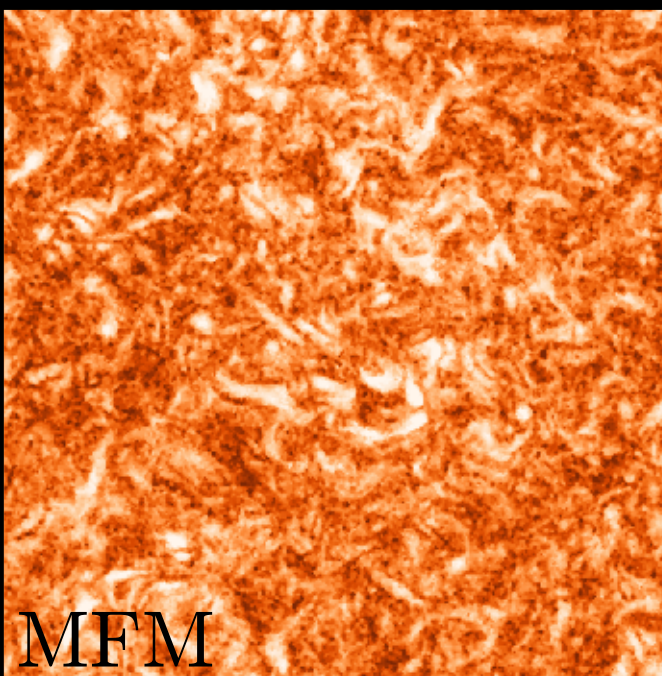
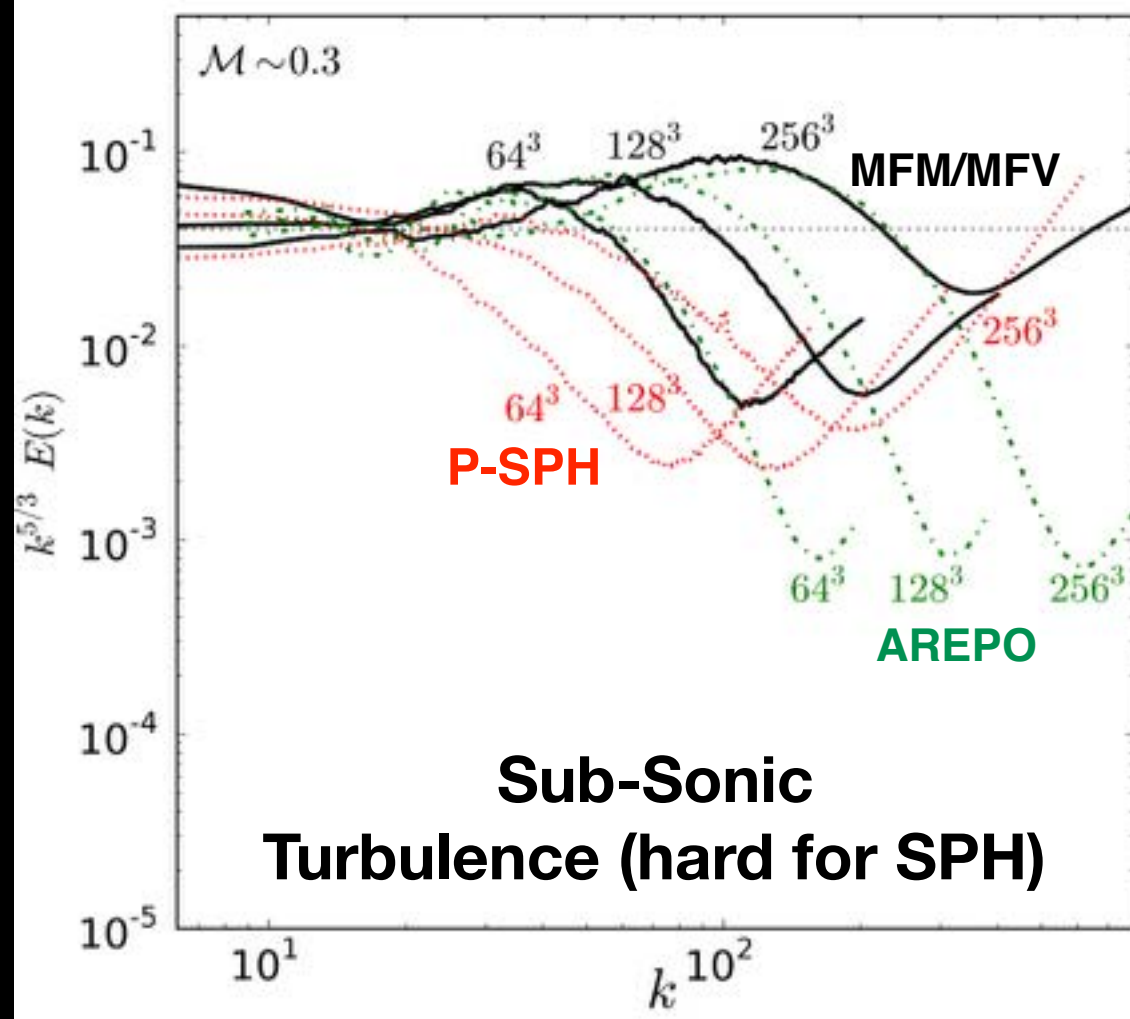
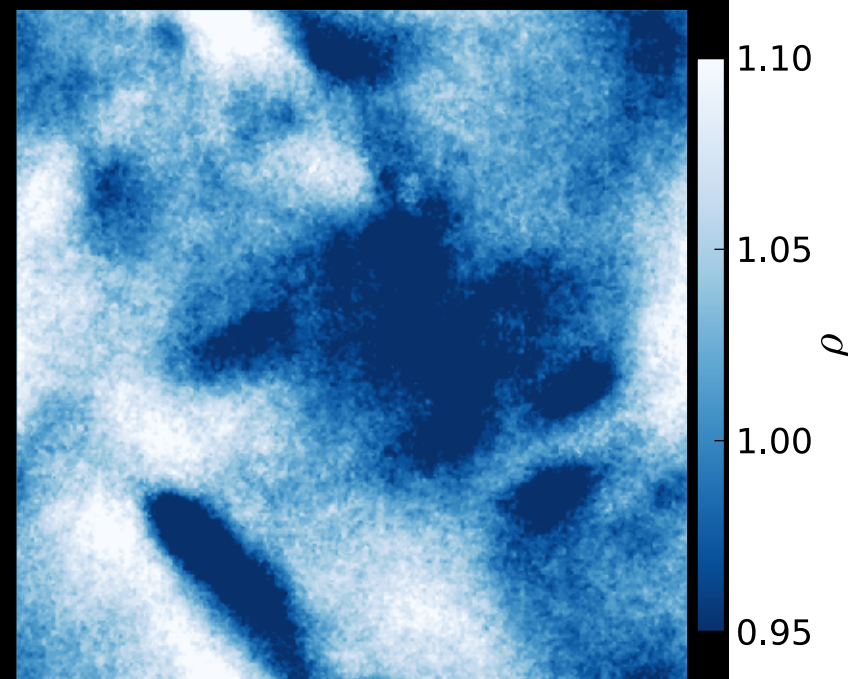
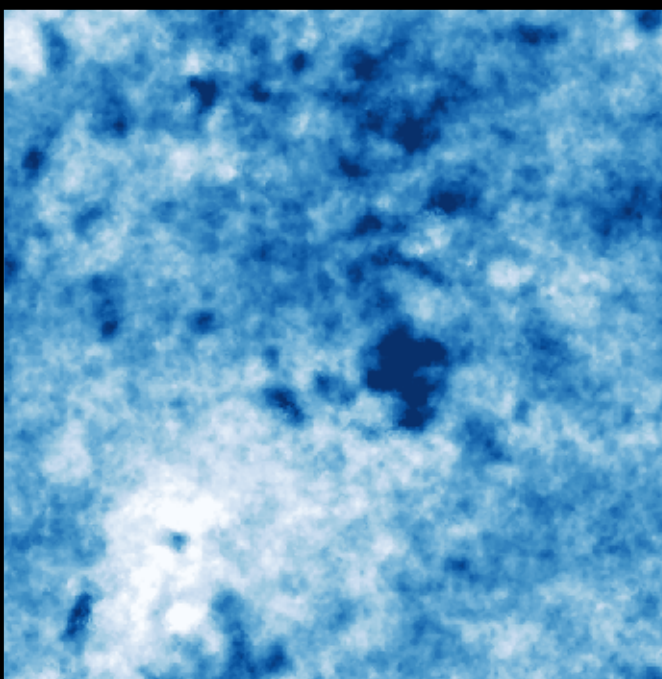
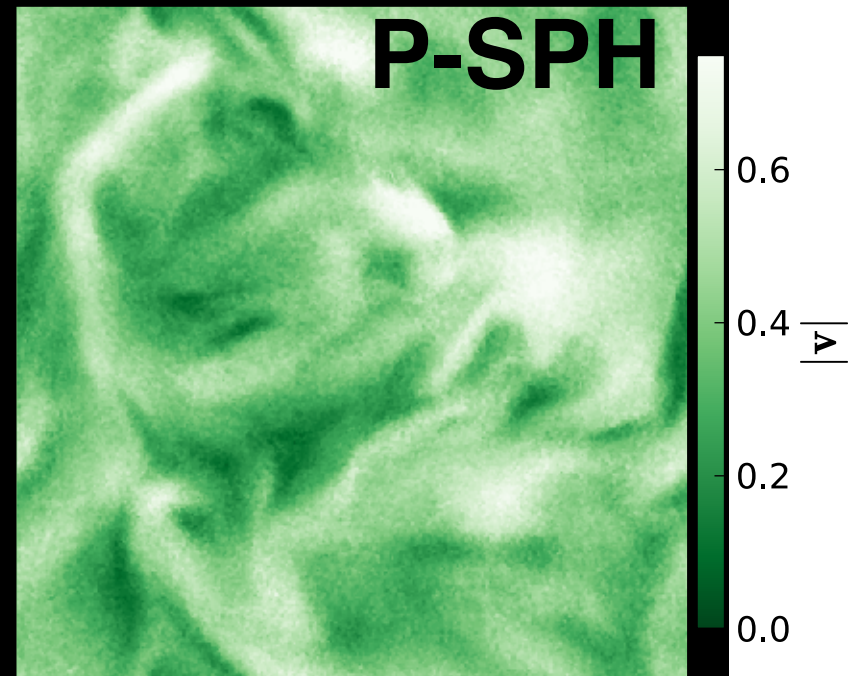
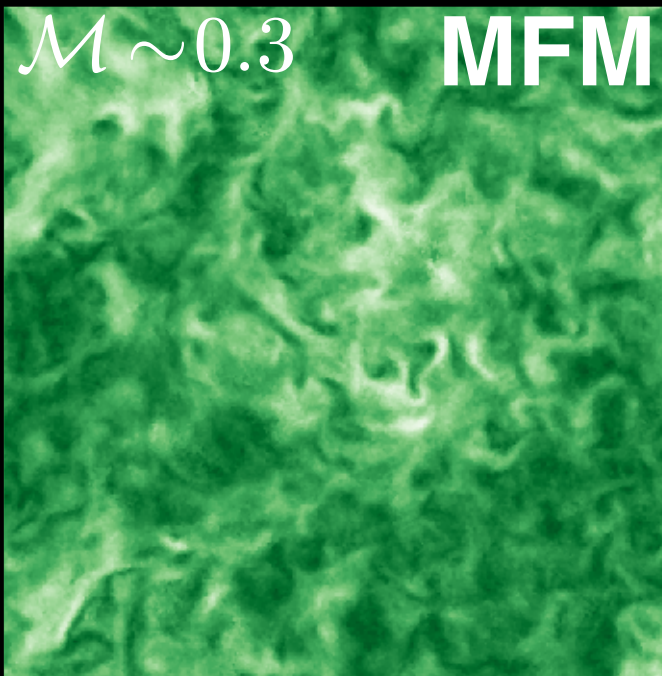
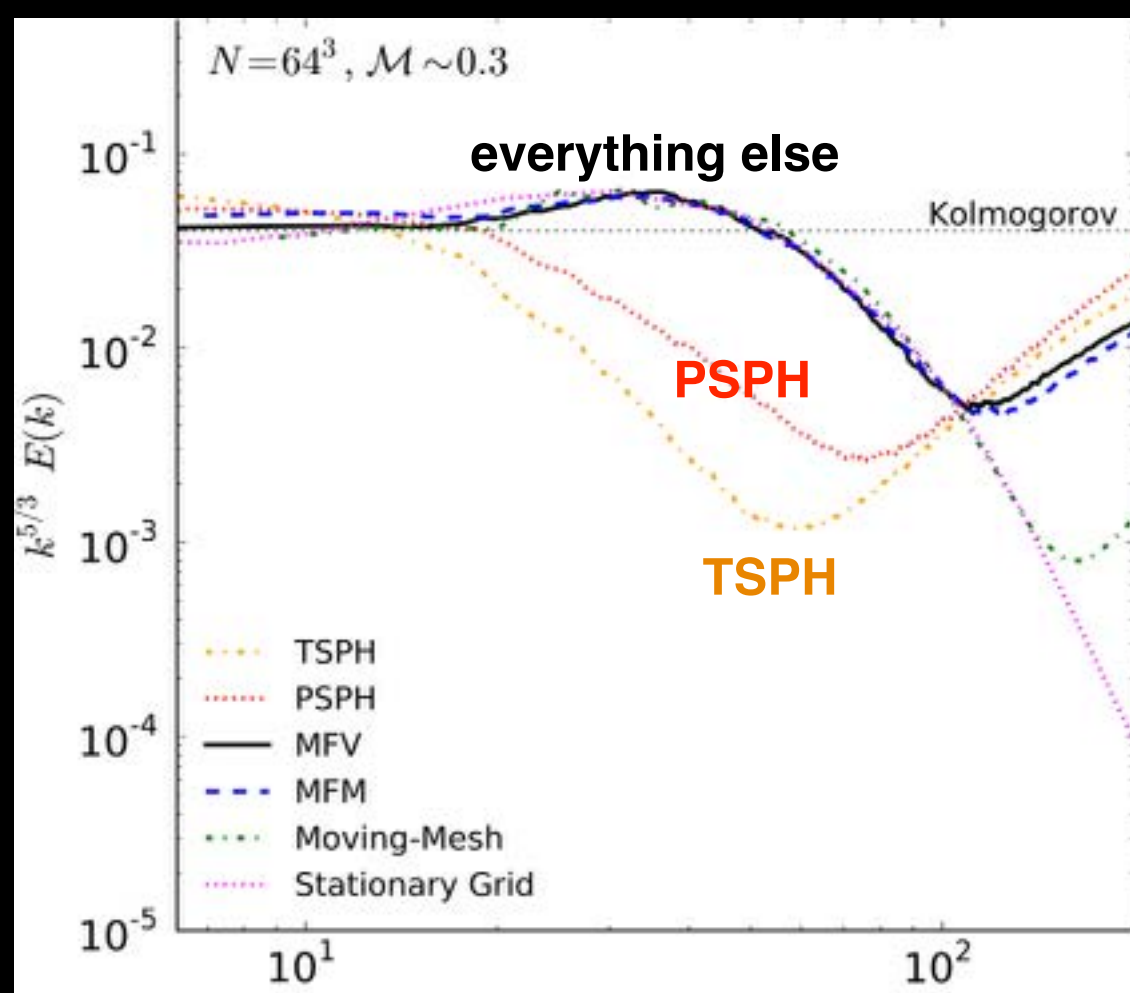
PSPH

Rayleigh-Taylor Instabilities

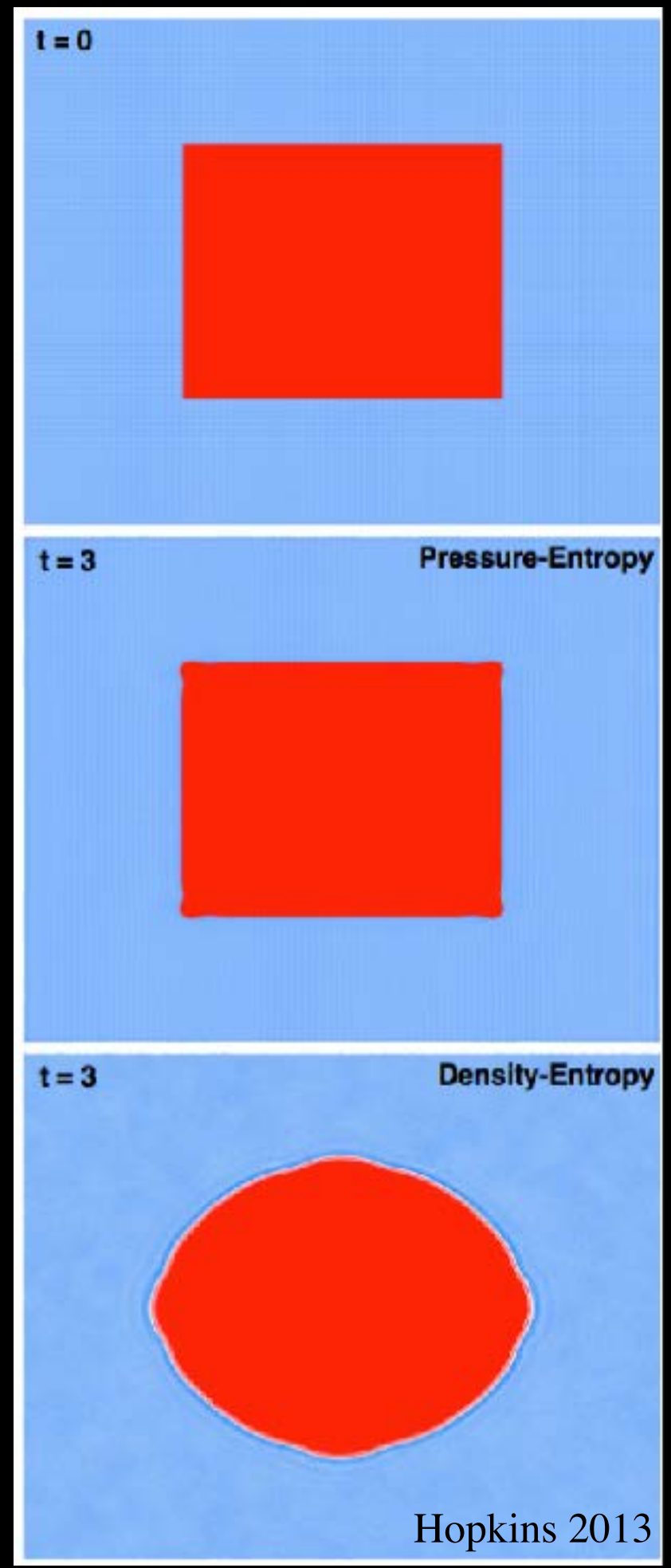
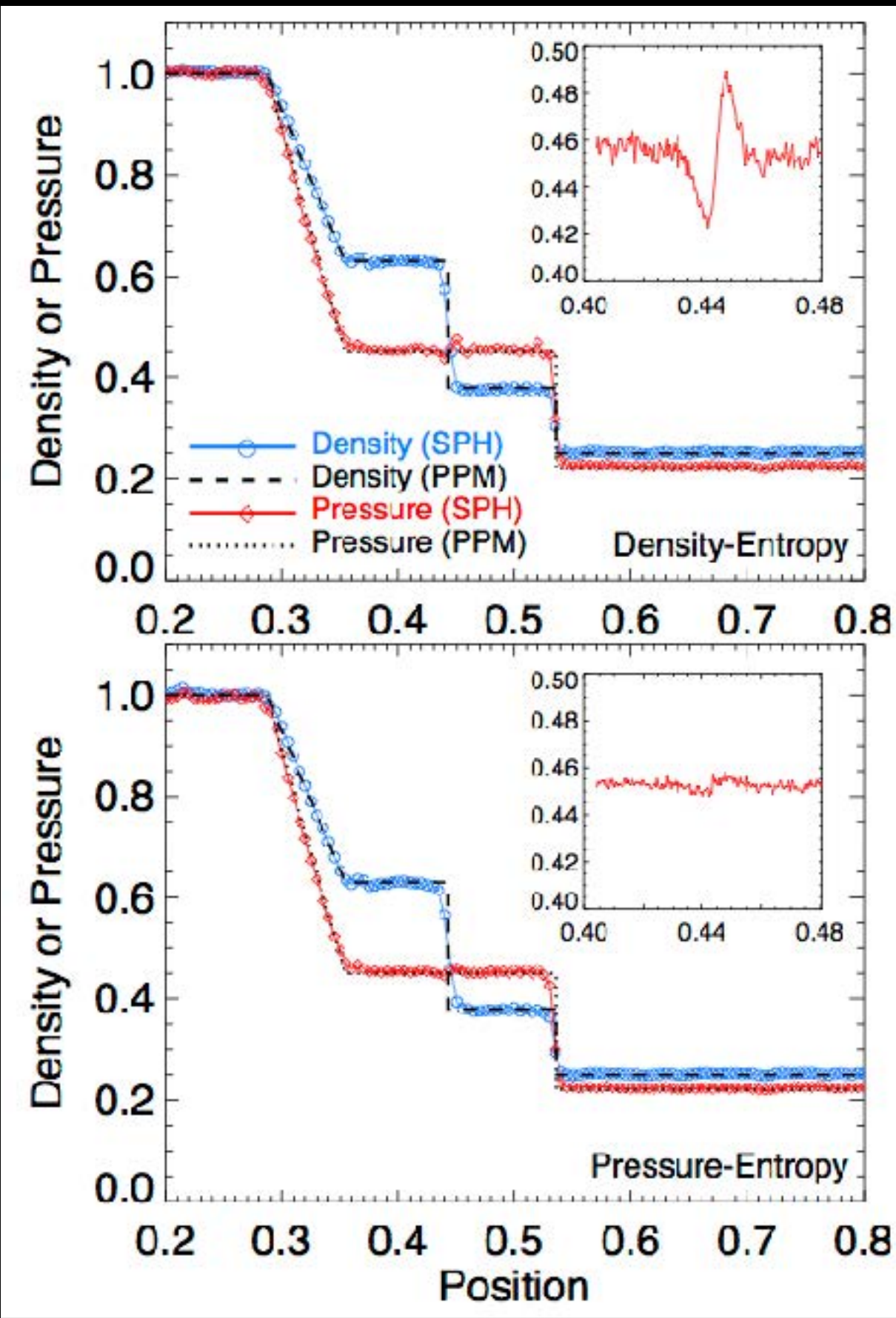


“The Blob” (Agertz et al.)





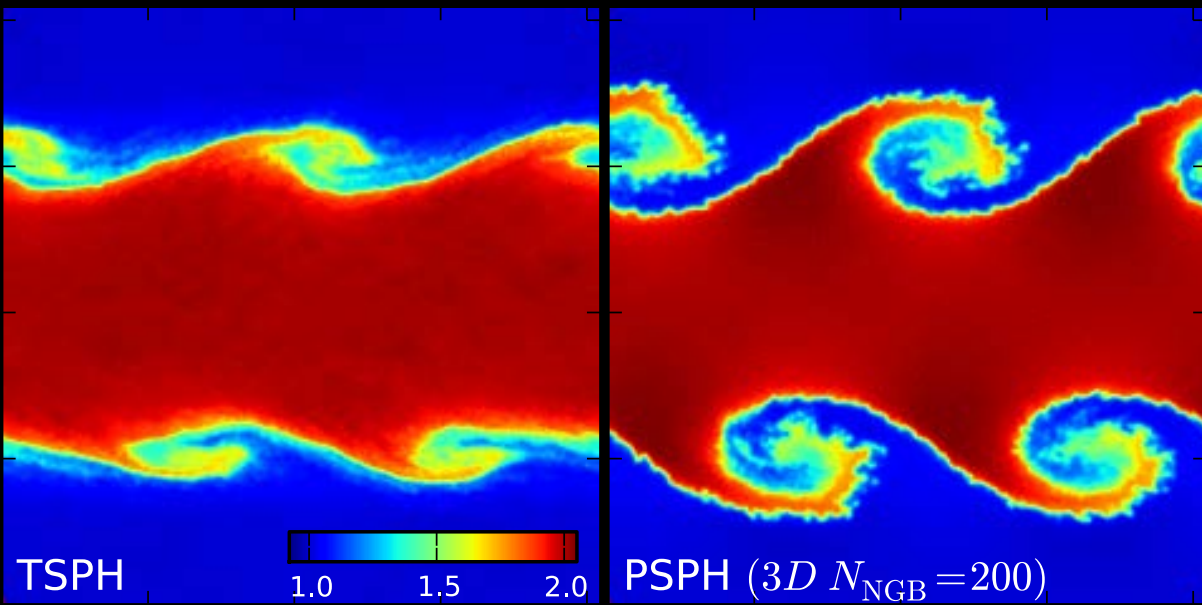
Artificial Surface Tension (entirely fix-able!)



Challenge:

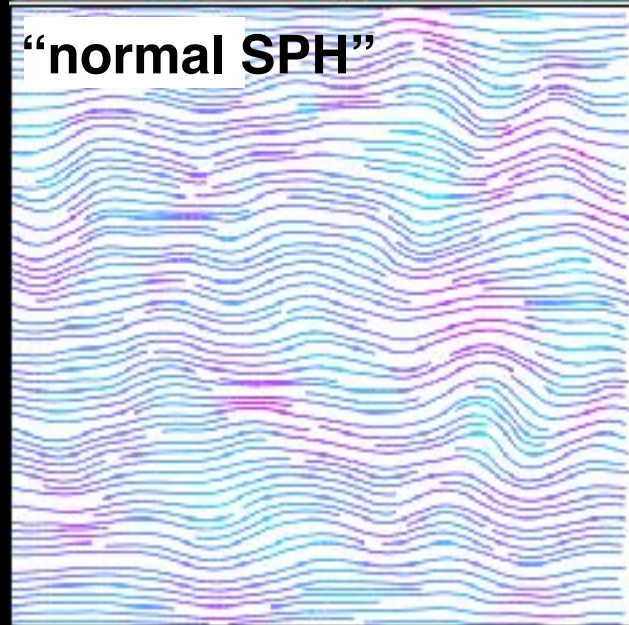
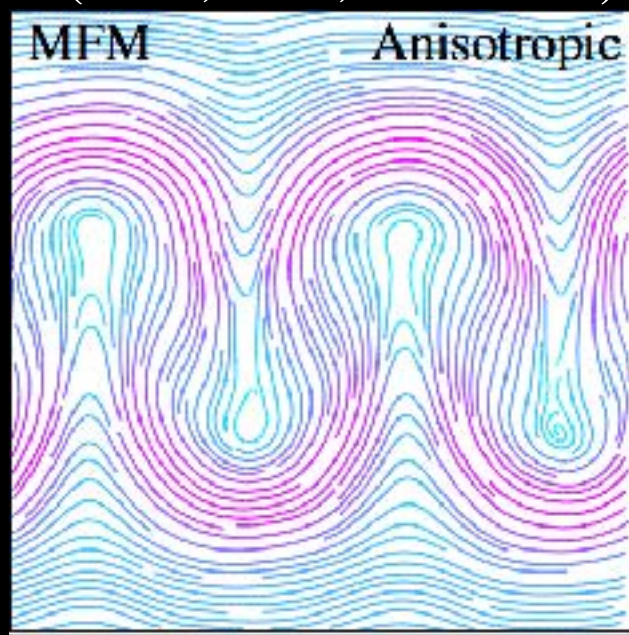
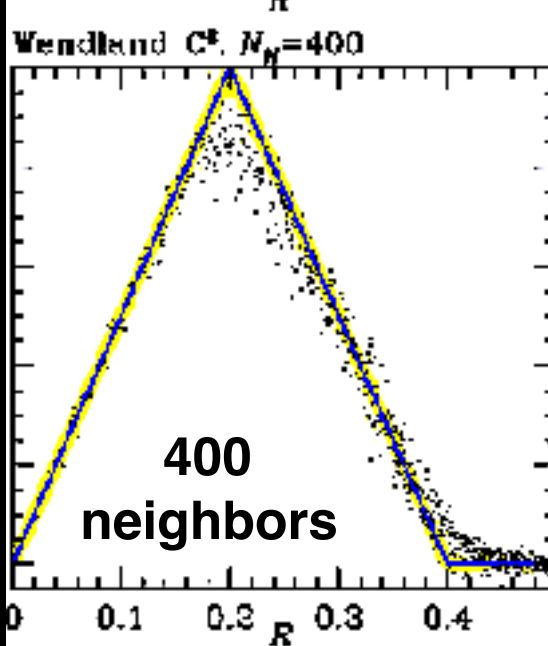
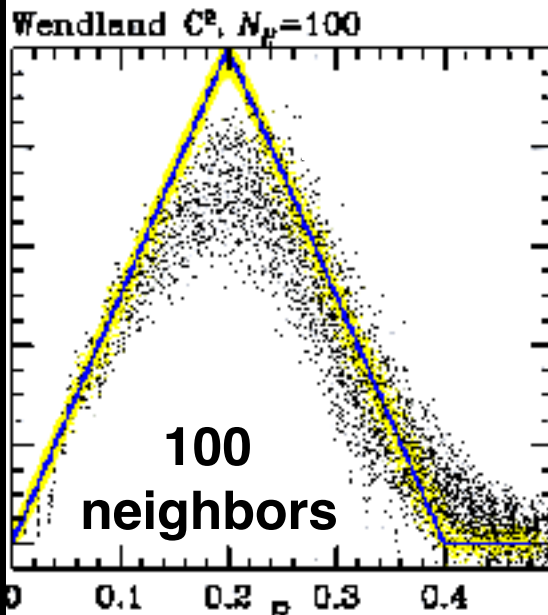
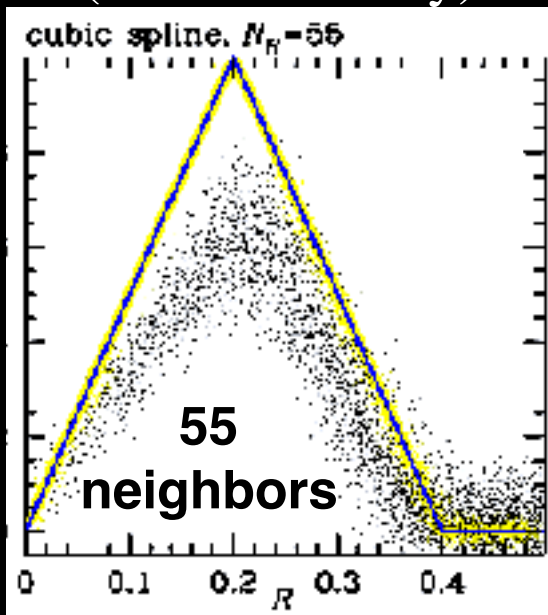
POPULAR METHODS FOR HYDRODYNAMICS HAVE PROBLEMS

- Fundamental low-order errors:
 - converge slowly:
 - “beat down” by increasing kernel size, but this is *not efficient!*
- MHD & anisotropic diffusion operators ill-posed



Gresho vortex (Dehnen & Aly)

Anisotropic Conduction (MTI, HBI, Hall MRI)

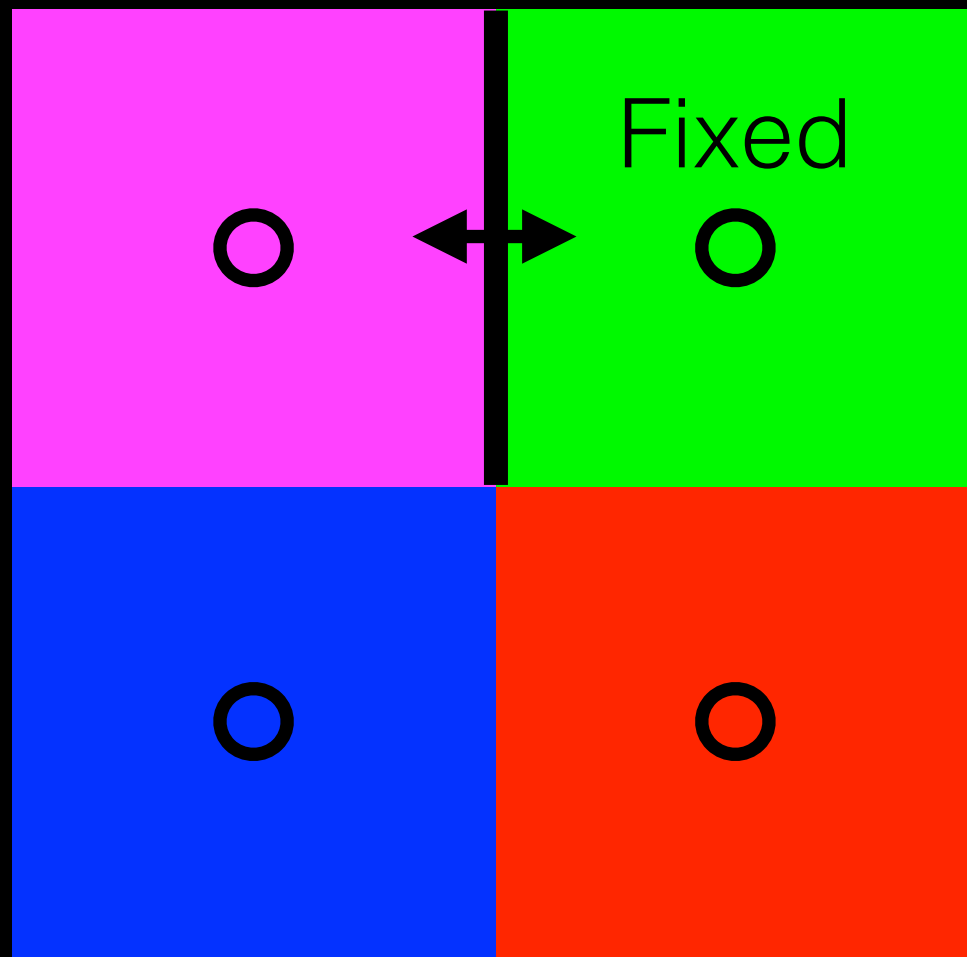


AMR + Fixed-Grids

Challenge:

POPULAR METHODS FOR
HYDRODYNAMICS HAVE PROBLEMS

Adaptive Mesh Refinement



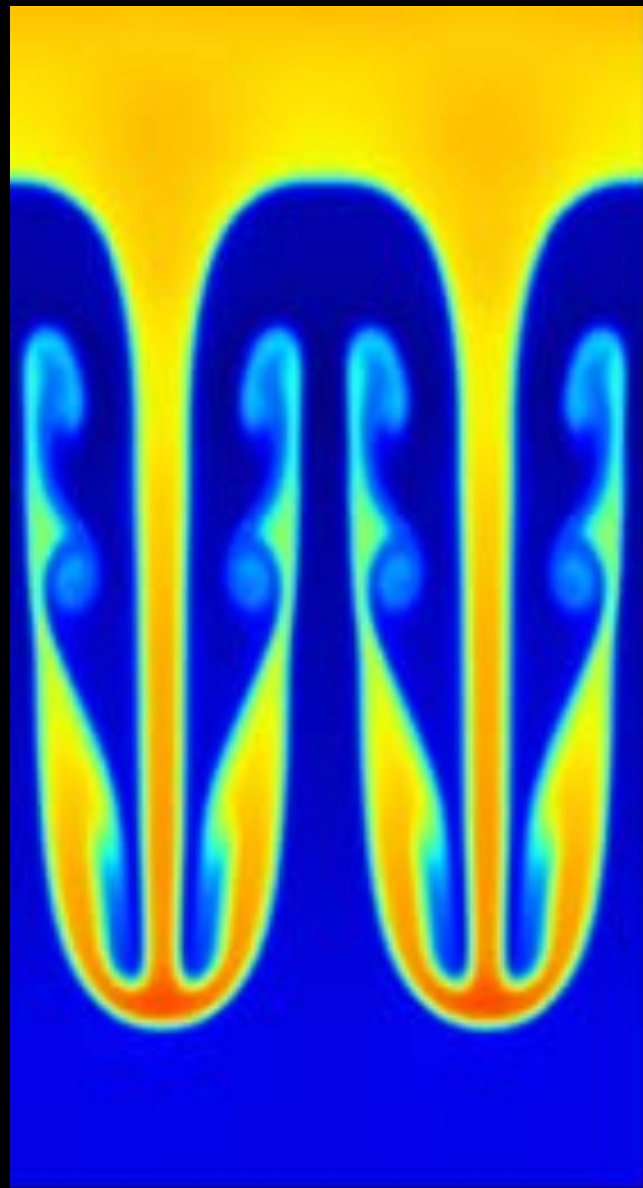
- Eulerian, well-studied, high-order
- Each cell carries conserved quantities inside volume V_i
- Solve Riemann problem between geometric faces

$$\Delta m_i = \int_{\text{cell}} \frac{\partial \rho}{\partial t} d^3 \mathbf{x} = - \int_{\text{cell}} \nabla \cdot (\rho \mathbf{v}) d^3 \mathbf{x}$$

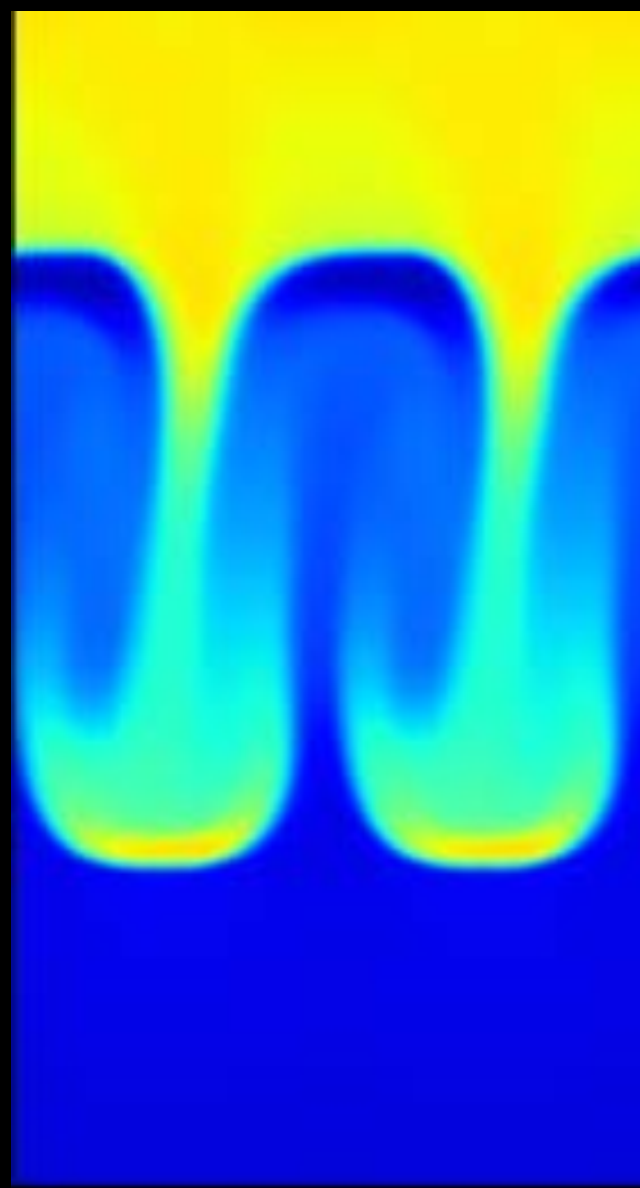
Adaptive Mesh Refinement (AMR)

CHALLENGE: POPULAR METHODS HAVE PROBLEMS

Rayleigh-Taylor instability
(AMR, 256^2)



(no bulk motion)

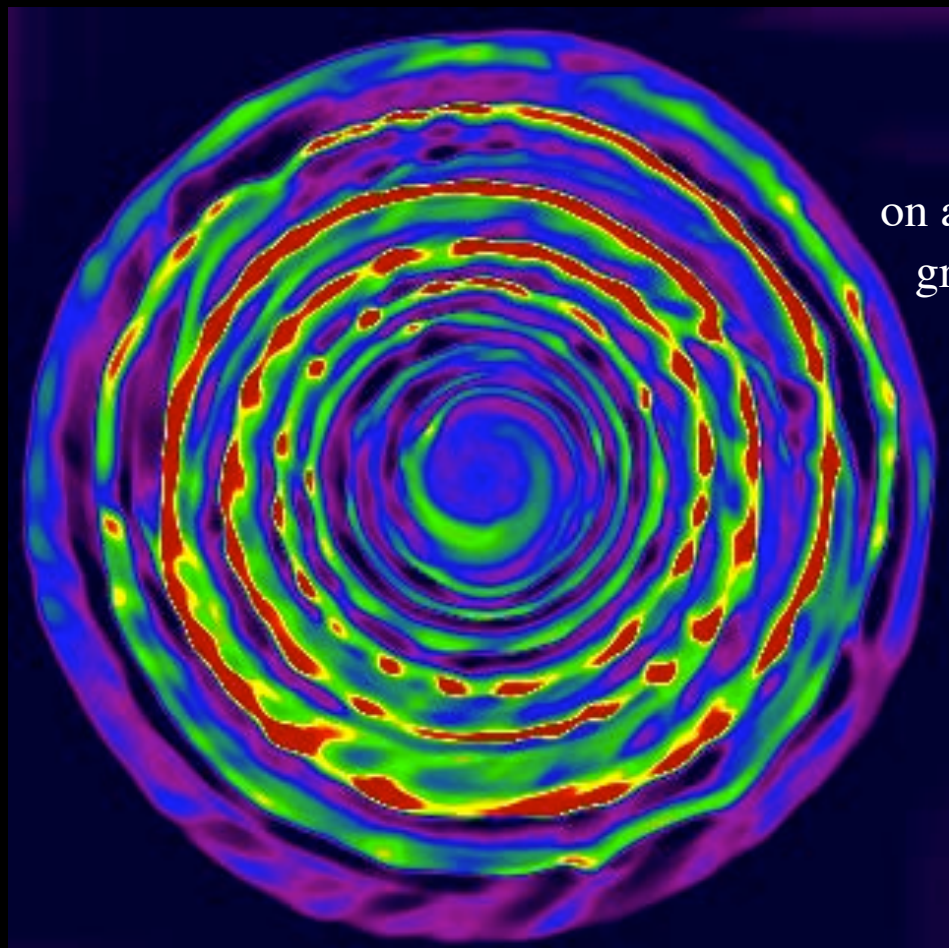


Mach 5 boost

- Eulerian, well-studied, high-order
- Excessive mixing/diffusion when fluid moves over cells

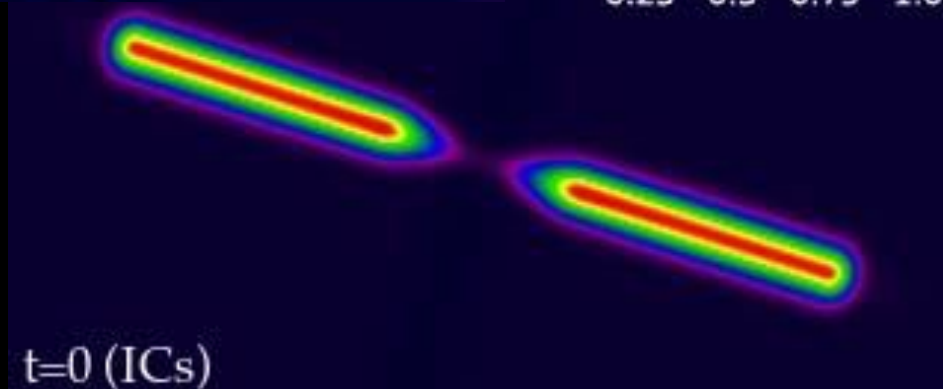
Adaptive Mesh Refinement (AMR)

CHALLENGE: POPULAR METHODS HAVE PROBLEMS

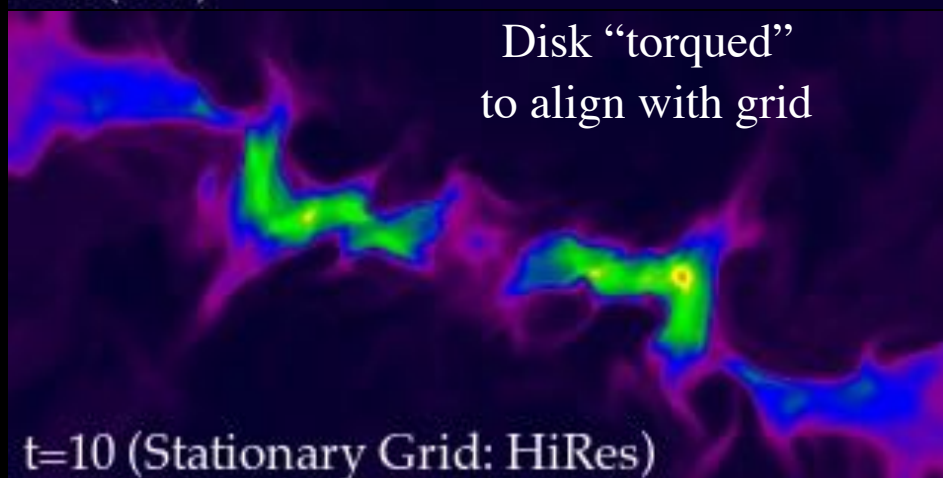


Keplerian disk
on a Cartesian (AMR)
grid after 10 orbits

0.25 0.5 0.75 1.0



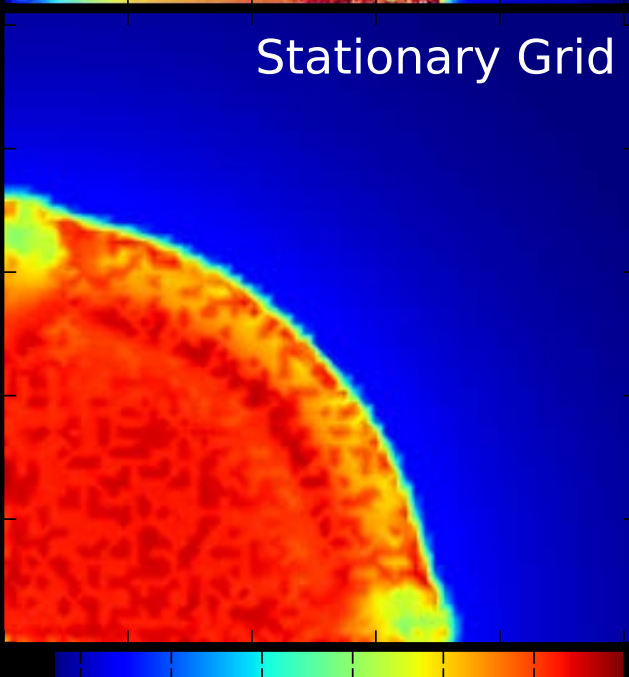
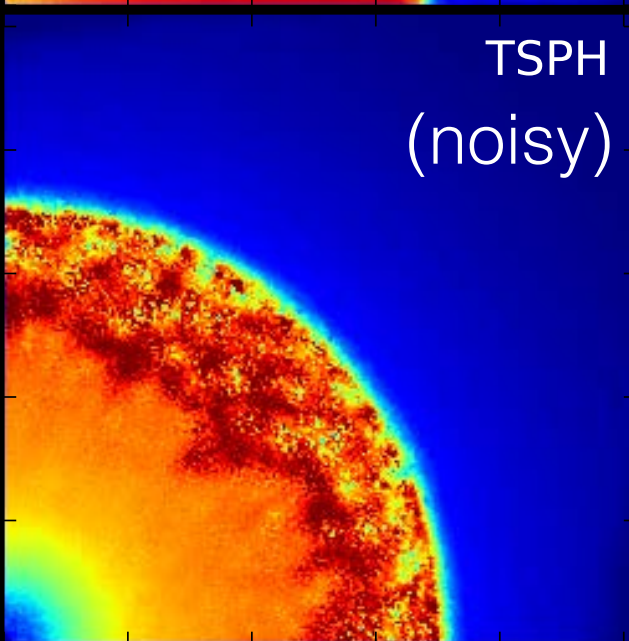
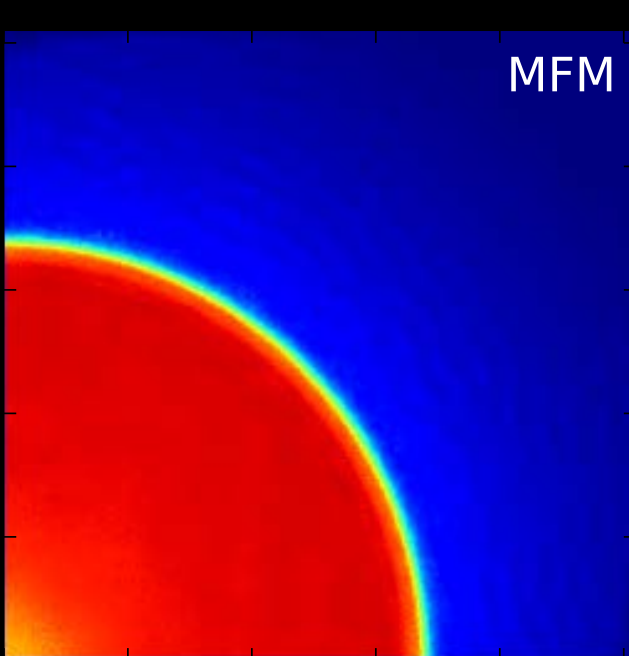
t=0 (ICs)



Disk "torqued"
to align with grid

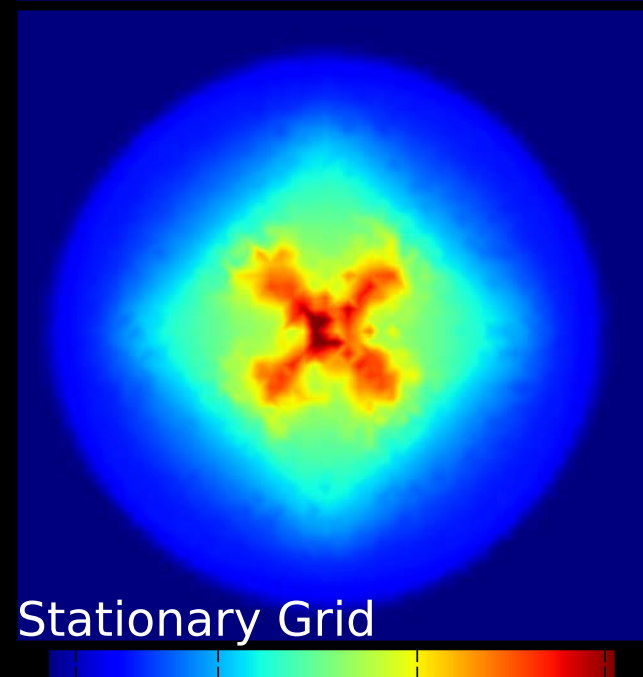
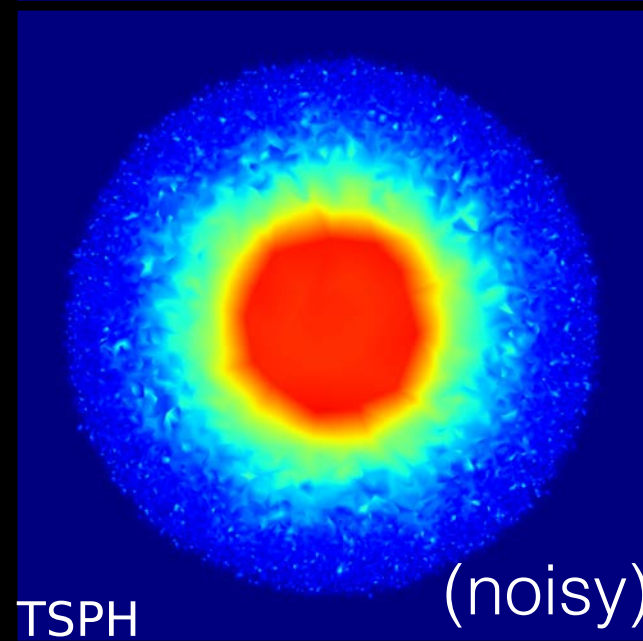
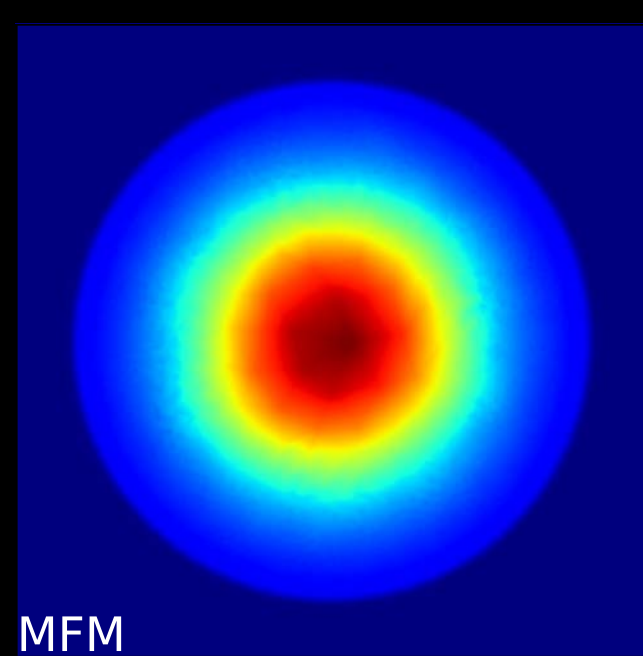
t=10 (Stationary Grid: HiRes)

- Eulerian, well-studied, high-order
- Excessive mixing/diffusion when fluid moves over cells
- Geometric effects:
 - carbuncle instability (shocks)
 - loss of angular momentum
 - grid-alignment (disks)
- Also "beaten down" with resolution, but *expensive*
 - Hahn '10: $\gg 512^2$ resolution to avoid grid-alignment



Noh
Implosion

Sedov
Explosion



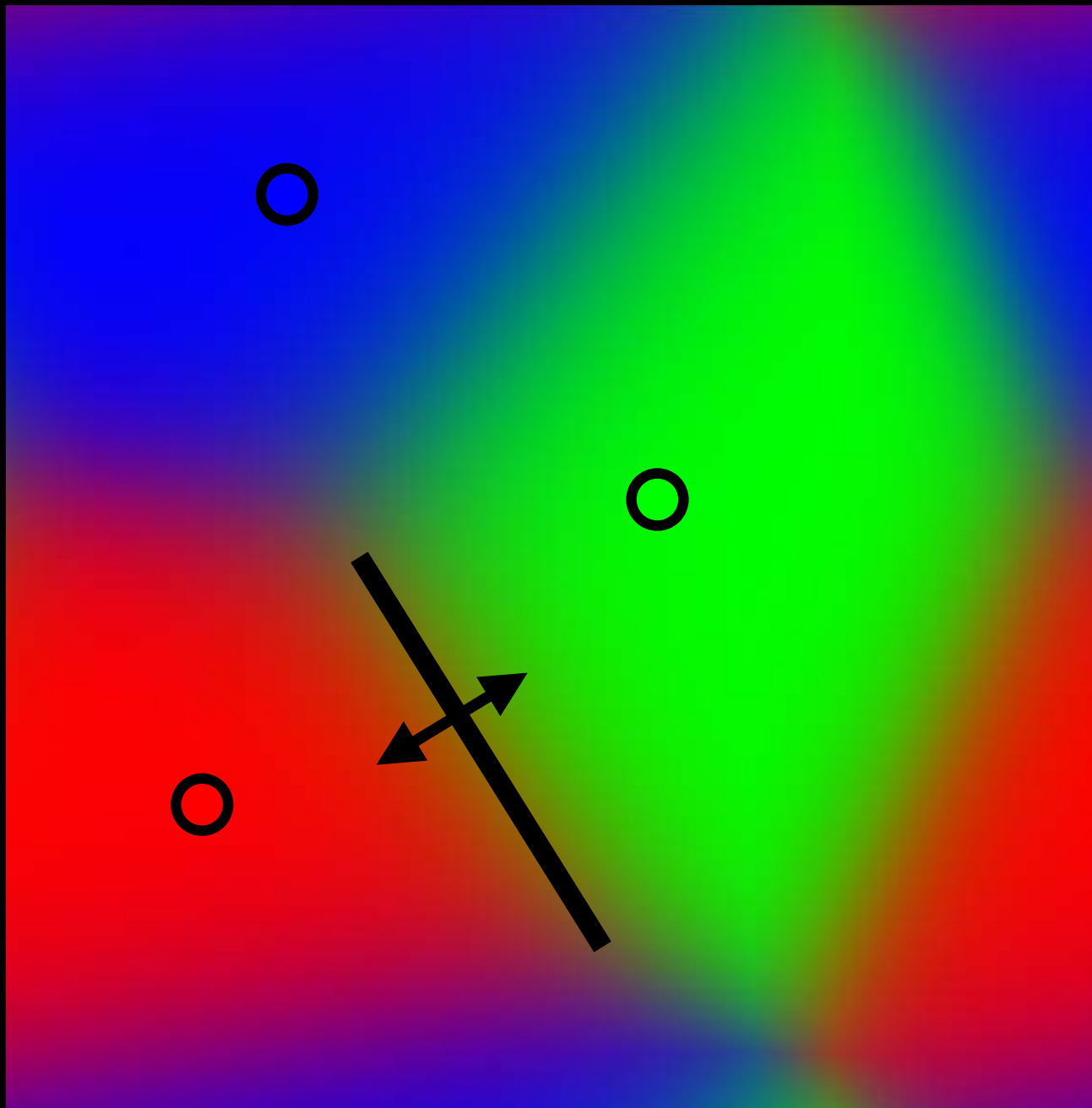
(grid breaks
what should be
spherical symmetry)



New Hybrid Lagrangian
Godunov Methods
(MFM/MFV/MMM)

Challenge:

POPULAR METHODS FOR
HYDRODYNAMICS HAVE PROBLEMS



New Methods Combine (some) Advantages of Both

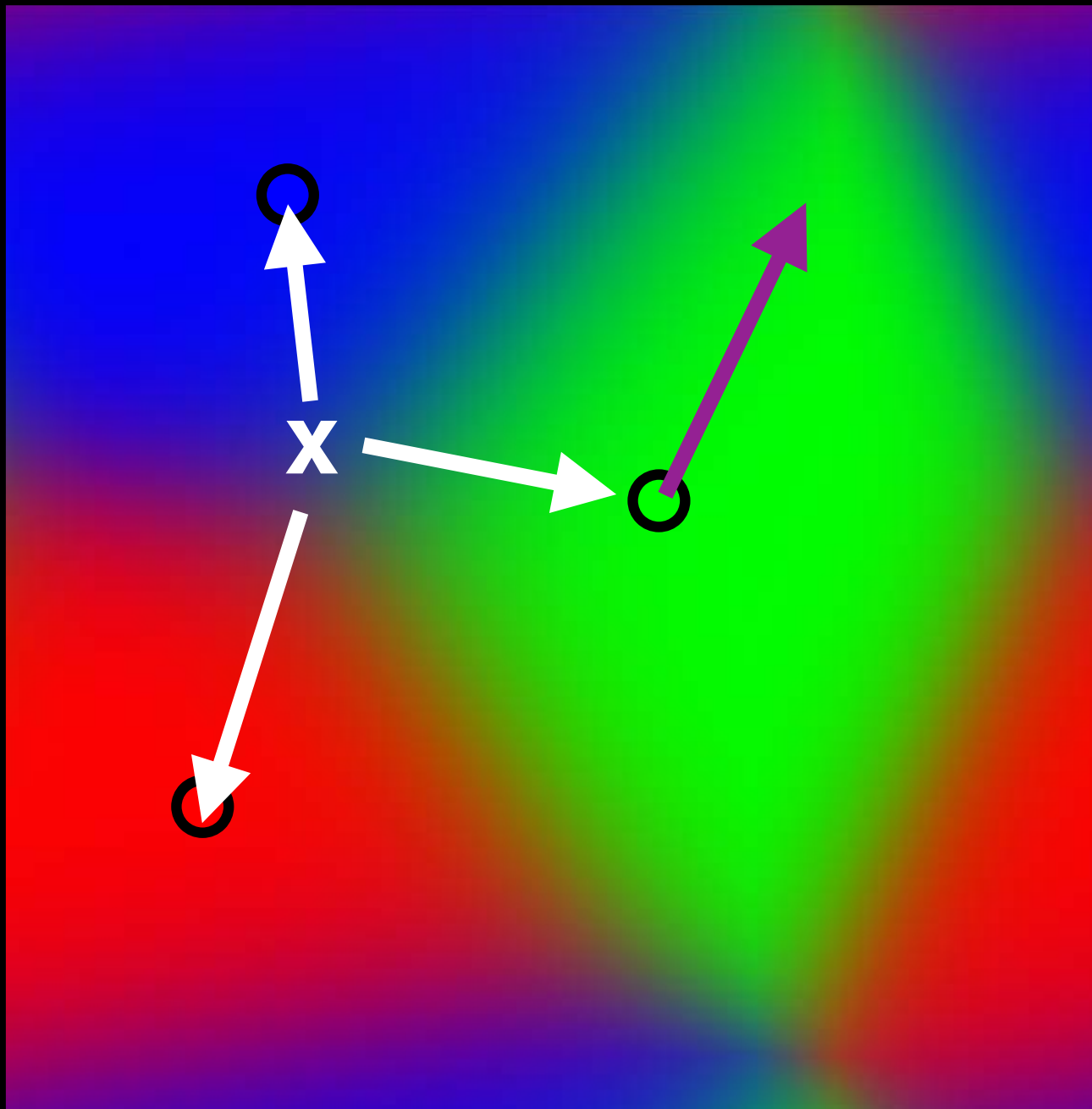
- Moving-meshes (AREPO), meshless finite-volume (GIZMO), high-order ALE methods
- Move with flow, no preferred geometry, but also accurate, high-order, and shock-capturing
- Less well-tested !

AREPO: Springel 2010

TESS/DISCO: Duffel 2011

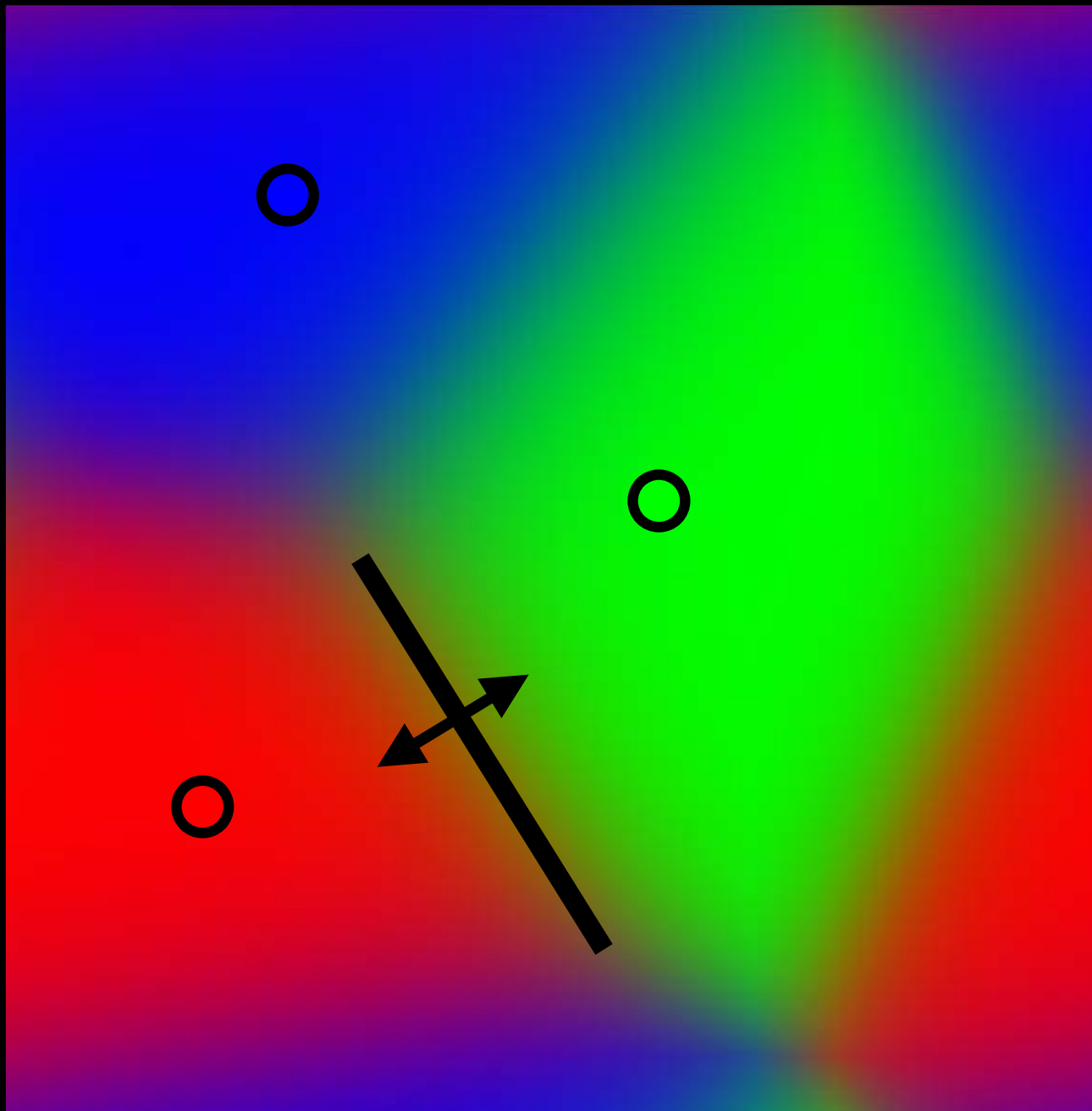
FVMHD3D: Gaburov 2012

GIZMO: Hopkins 2015 (arXiv:1409.7395)

Challenge:POPULAR METHODS FOR
HYDRODYNAMICS HAVE PROBLEMS

- Mesh-generating points move (if desired)
- Volume is “partitioned” with a continuous kernel (MFM/MFV) or step function (moving-mesh)

$$d\text{Vol}_{i,j,k} = d^3\mathbf{x} \frac{W(\mathbf{x} - \mathbf{x}_{i,j,k})}{\sum W_{i,j,k}}$$

Challenge:POPULAR METHODS FOR
HYDRODYNAMICS HAVE PROBLEMS

- Integrate EOM over volume:
equivalent to Riemann problem
at “effective face” (quadrature)

$$\Delta m_i = \int_{\text{vol}} \frac{\partial \rho}{\partial t} d^3 \mathbf{x} = - \int_{\text{vol}} \nabla \cdot (\rho \mathbf{v}) d^3 \mathbf{x}$$

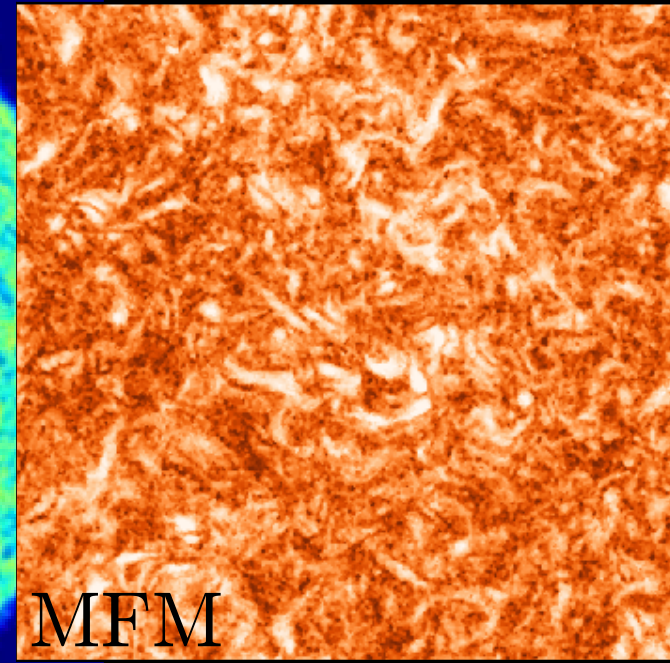
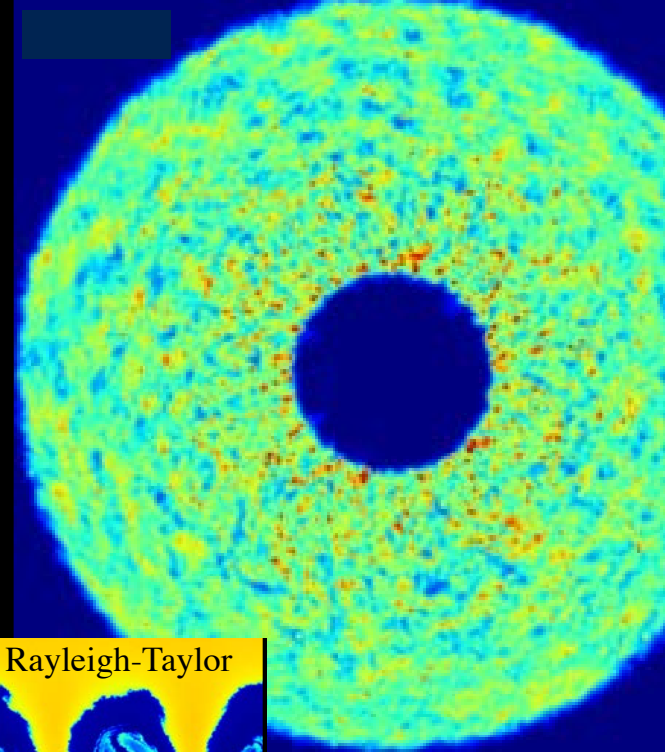
New Methods Combine (some) Advantages of Both:

(BUT REMAIN LESS WELL-TESTED)

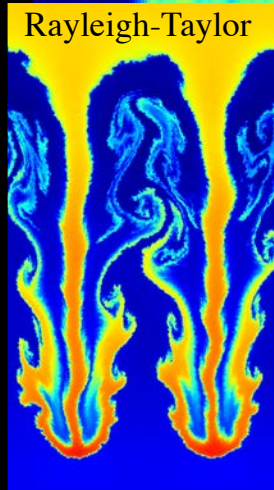
- Moving-meshes (AREPO), meshless finite-volume (GIZMO), high-order ALE methods
- Move with flow, no preferred geometry, but also accurate, high-order, and shock-capturing
- Grid noise is more severe

GIZMO: disk after 100 orbits

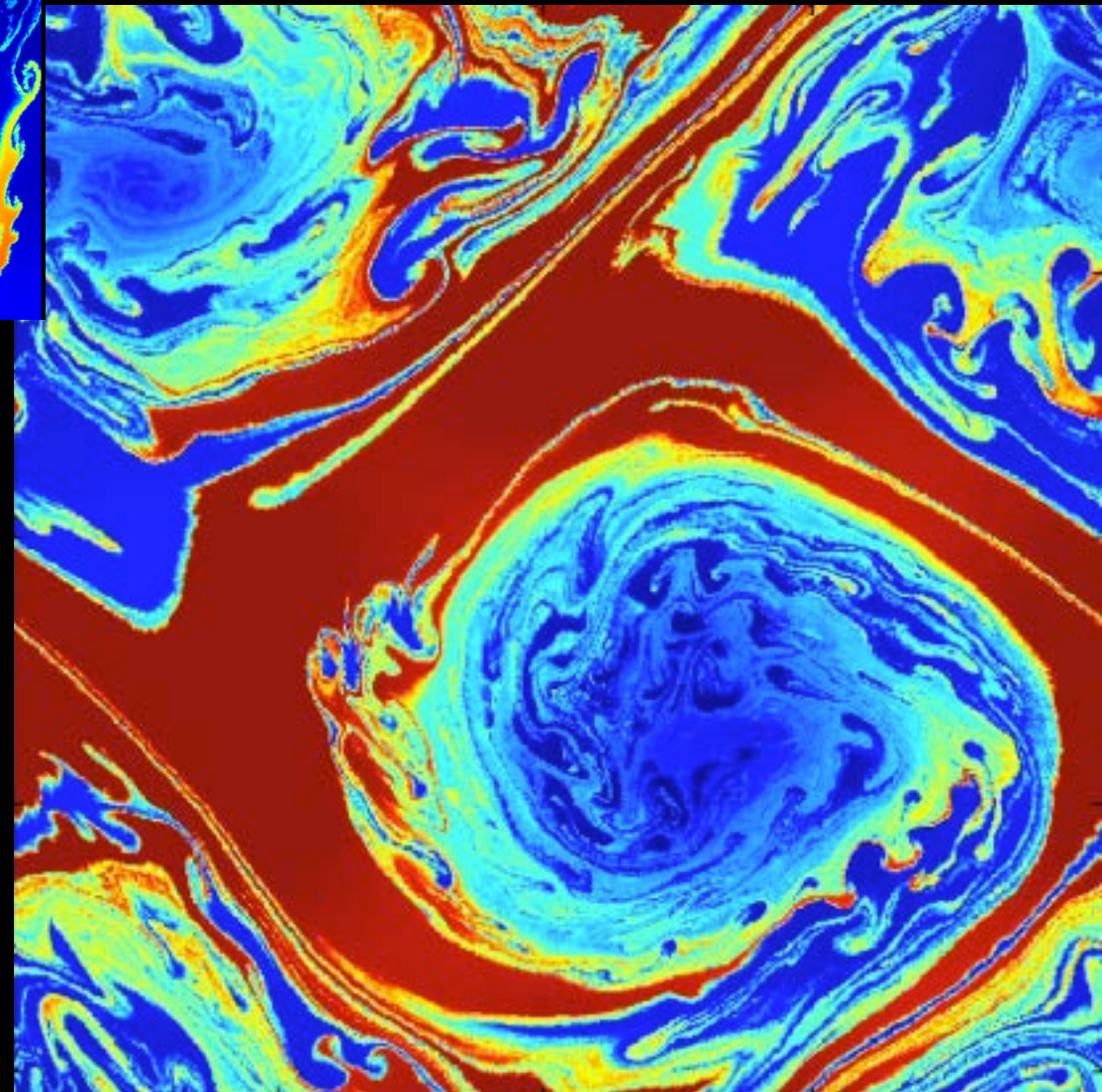
sub-sonic turbulence



MFM



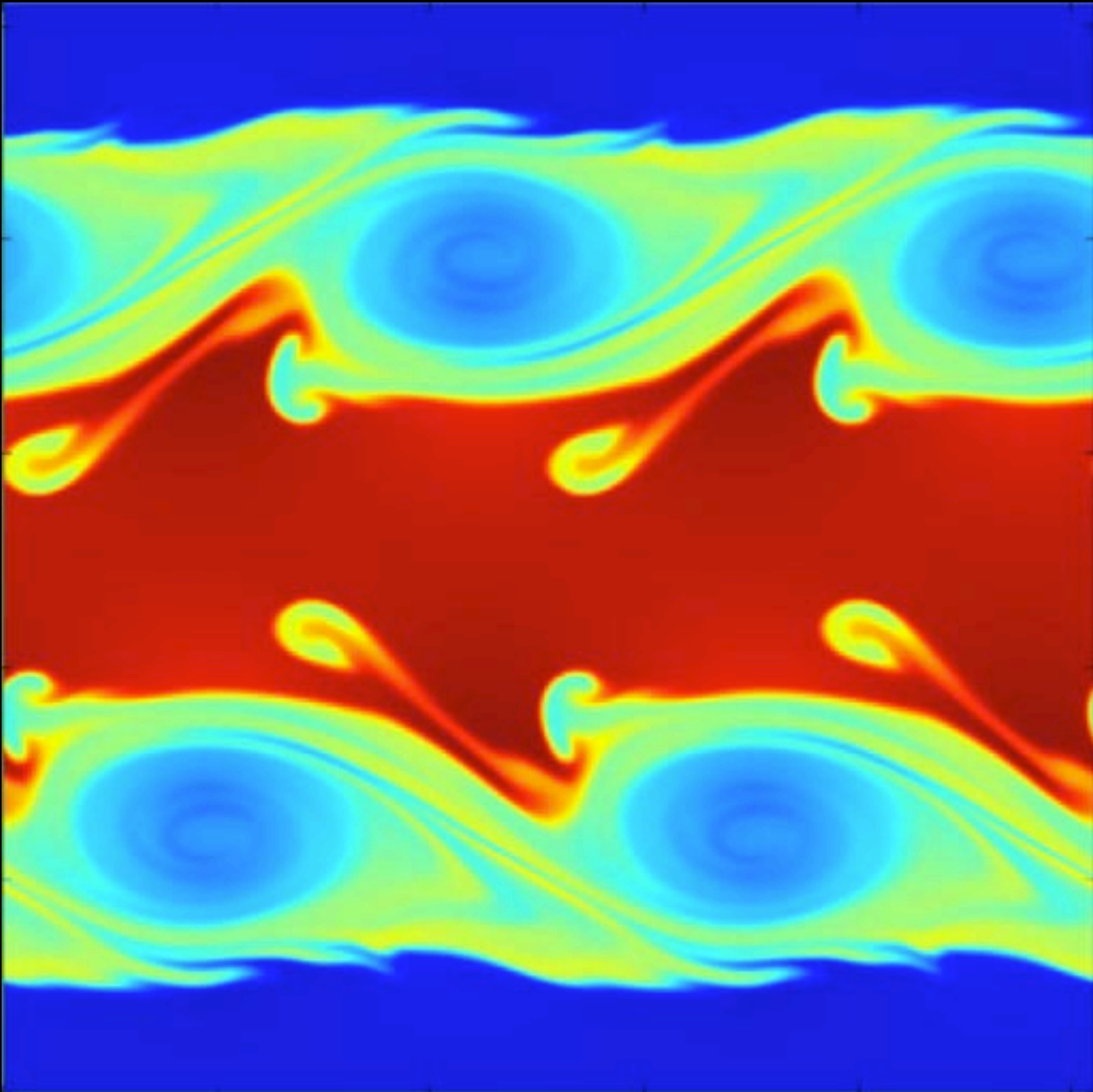
Rayleigh-Taylor



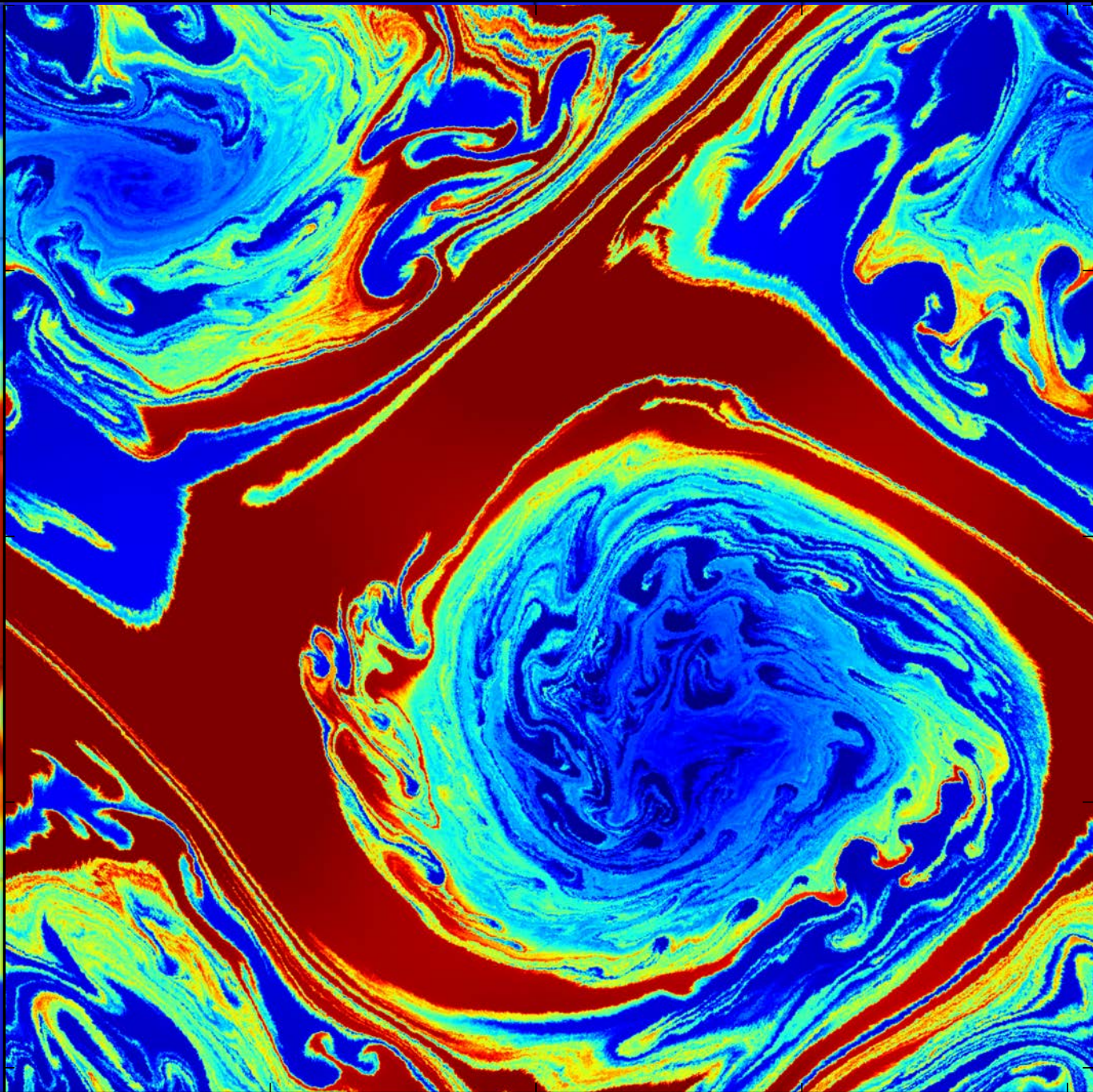
AREPO: Springel 2010
TESS/DISCO: Duffel 2011
FVMHD3D: Gaburov 2012
GIZMO: Hopkins 2015

GIZMO: New Meshless Methods & Fluid Mixing

(www.tapir.caltech.edu/~phopkins)



Cartesian Grid



Meshless Finite Volume

Summary

SPH:

Neighbor number: is it worth it?

Hot halos: numerical mixing

MHD: need *even more* neighbors

Anisotropic diffusion: fundamental barriers

Moving-mesh, meshless Godunov:

“Grid noise”: Mach $< \sim 0.01$ problems

MHD: div-cleaning corrupts

weak-field ($v_A < 0.01 * v_{\text{turb,thermal}}$)

AMR:

Angular momentum/grid alignment